



# Geological Setting of the Luna Glob Candidate Landing Site Region in the Northern High Latitudes: Characterization from LOLA and Related Data

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Landing Site Selection for LUNA-GLOB mission  
International Workshop #2

*Moscow, Institute for Space Research (IKI), May 31 – June 2, 2011*

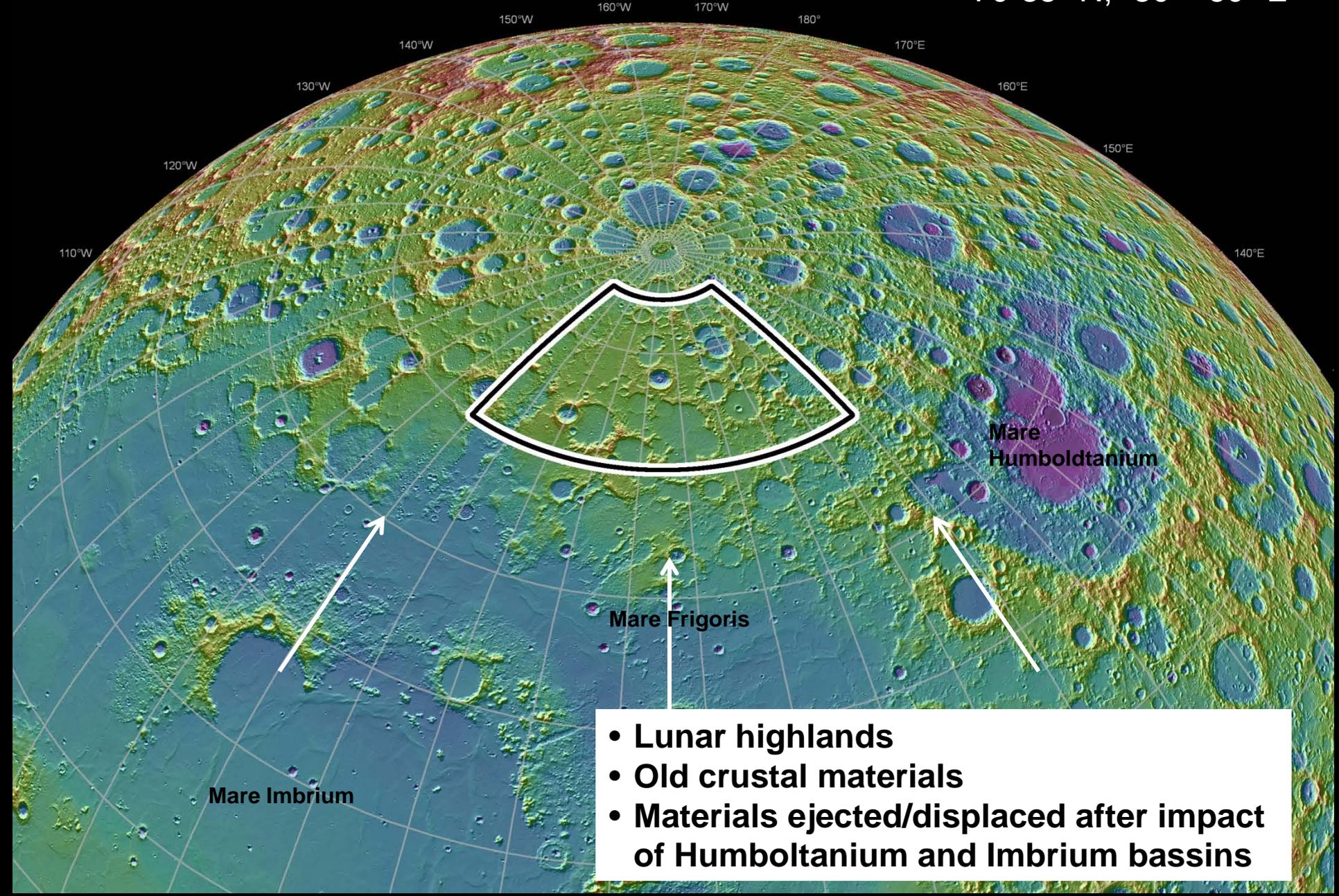


# Criteria of the Luna Glob Candidate Landing Site selection:

1. NPOL Ballistic constraints: The boundaries of landing area:  
 $70-85^{\circ}$  N,  $-30 - 60^{\circ}$  E
2. High Volatile Enrichment (by LEND data, Low neutron flux).
3. Safety of approaching trace (LOLA data).
4. The landing area (15x30 km ellipse) should be safely smooth (LROC, LOLA, Micro-RF data).

# Luna Glob Site Region

NPOL: The boundaries of landing area:  
70-85° N, -30 – 60° E

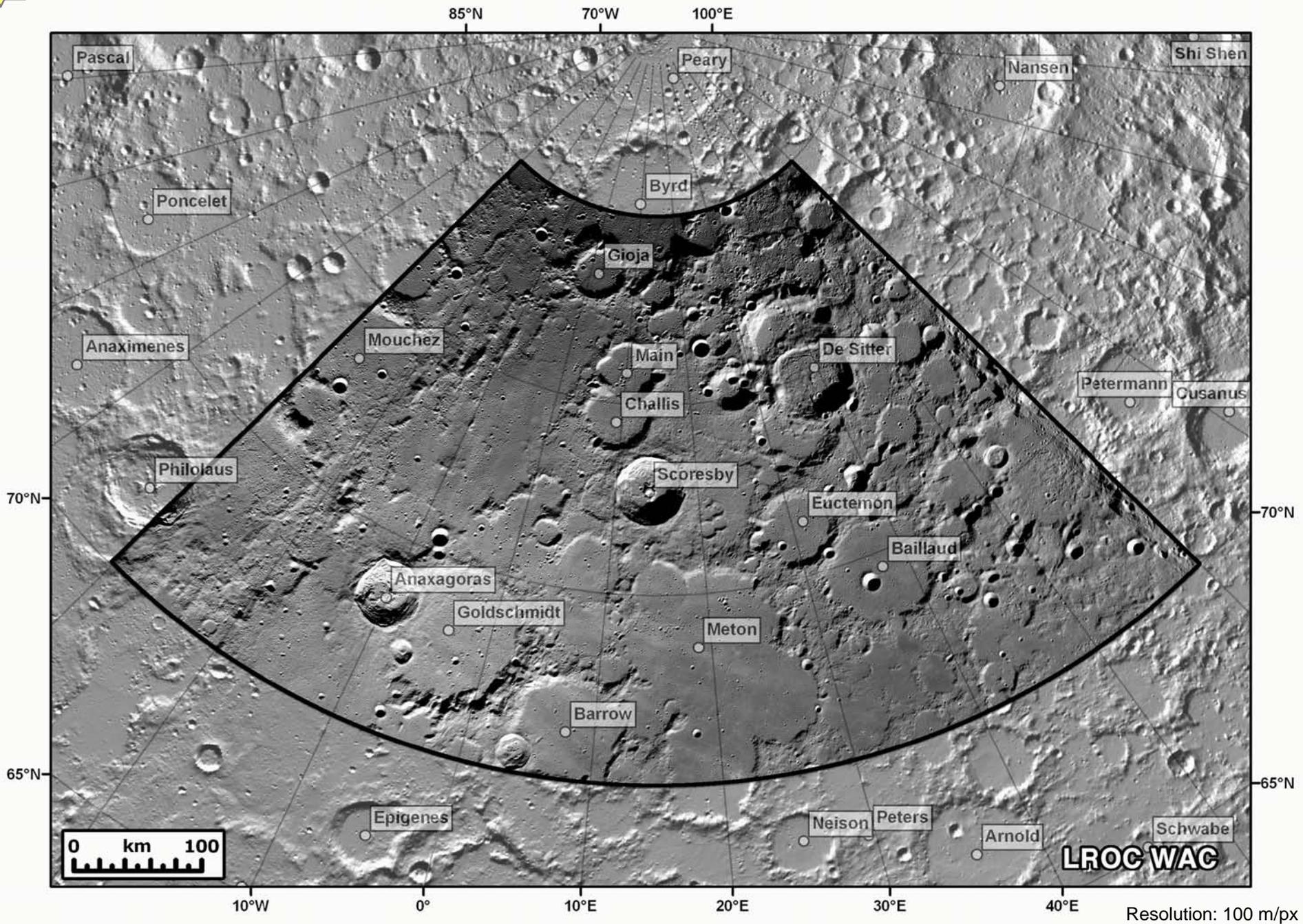


Mare Imbrium

Mare Frigoris

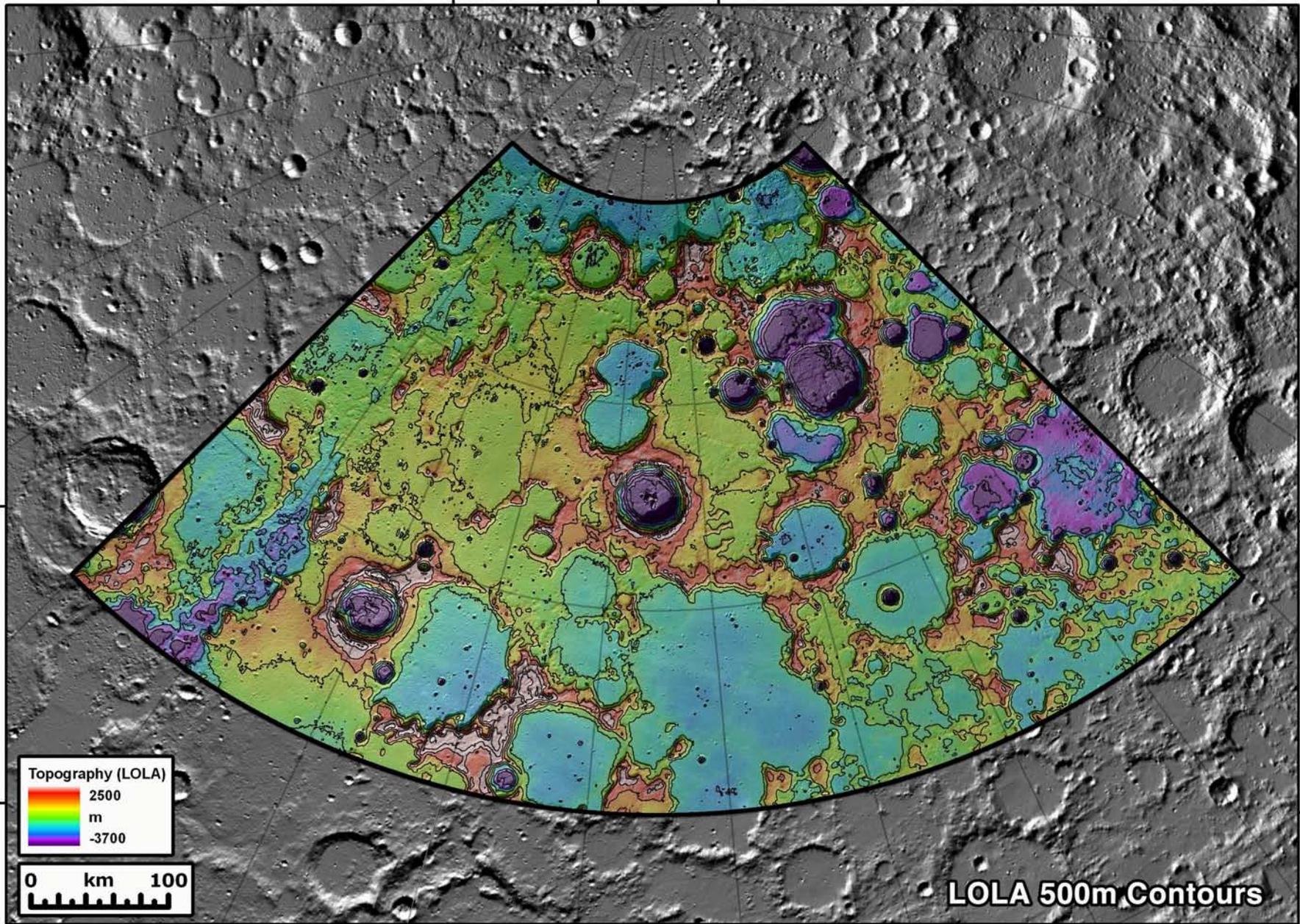
Mare Humboldtianum

- Lunar highlands
- Old crustal materials
- Materials ejected/displaced after impact of Humboldtianum and Imbrium basins





85°N 70°W 100°E



Topography (LOLA)  
2500  
m  
-3700

0 km 100

LOLA 500m Contours

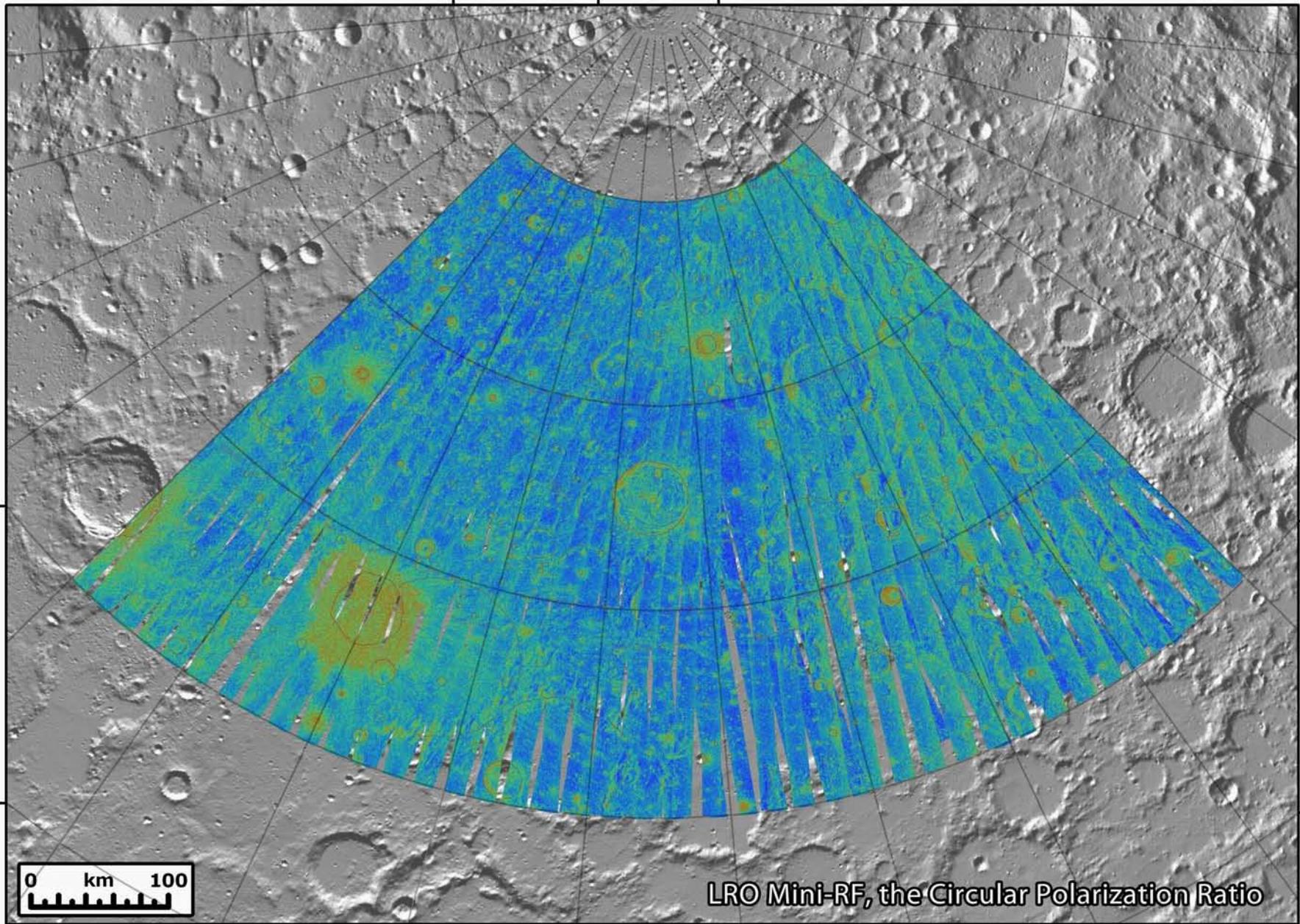
70°N  
65°N

70°N  
65°N

10°W 0° 10°E 20°E 30°E 40°E



85°N 70°W 100°E



70°N

70°N

65°N

65°N

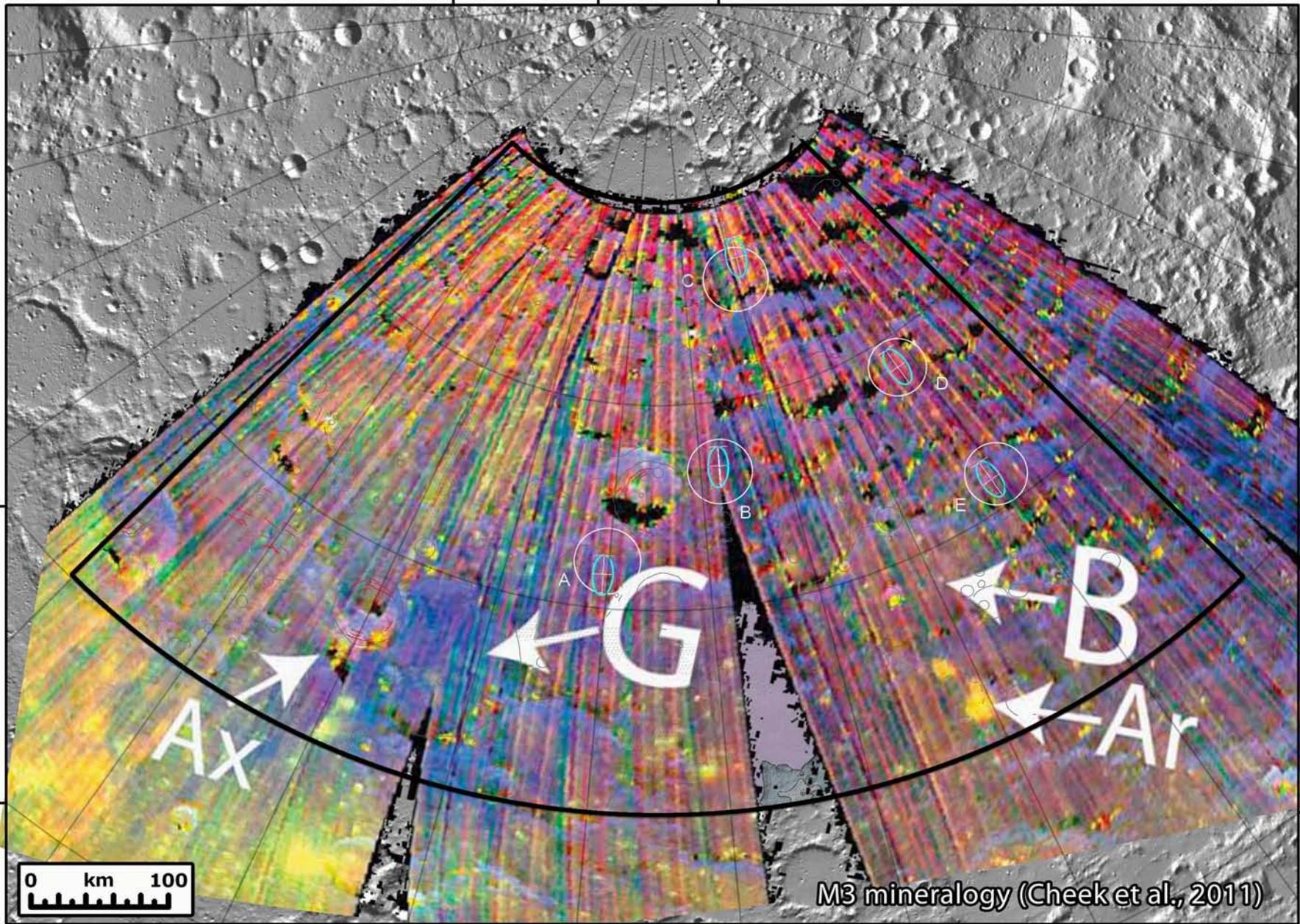


LRO Mini-RF, the Circular Polarization Ratio

10°W 0° 10°E 20°E 30°E 40°E



85°N 70°W 100°E



70°N

70°N

65°N

65°N

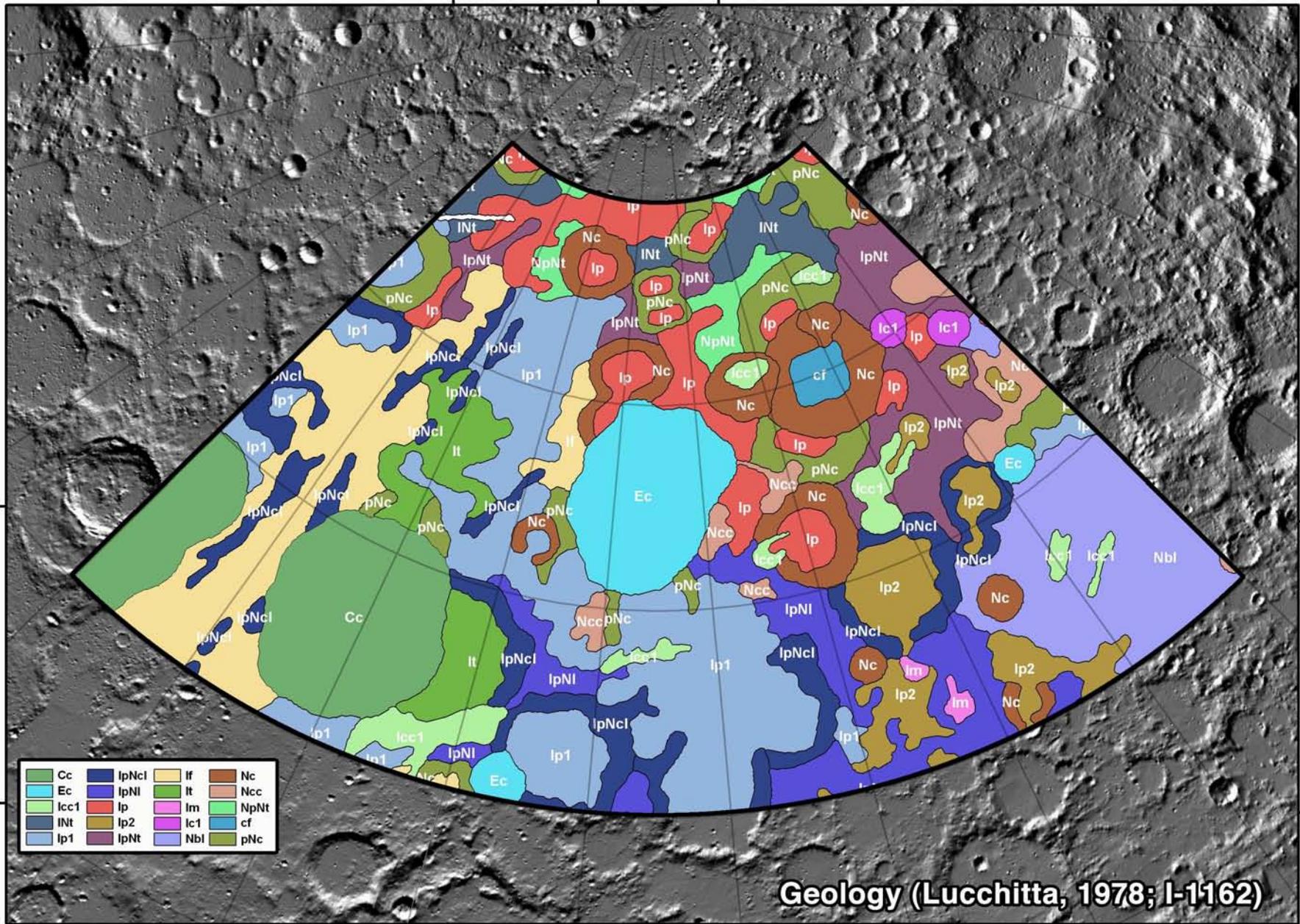
0 km 100

M3 mineralogy (Cheek et al., 2011)

10°W 0° 10°E 20°E 30°E 40°E



85°N 70°W 100°E



70°N

70°N

65°N

65°N

10°W

0°

10°E

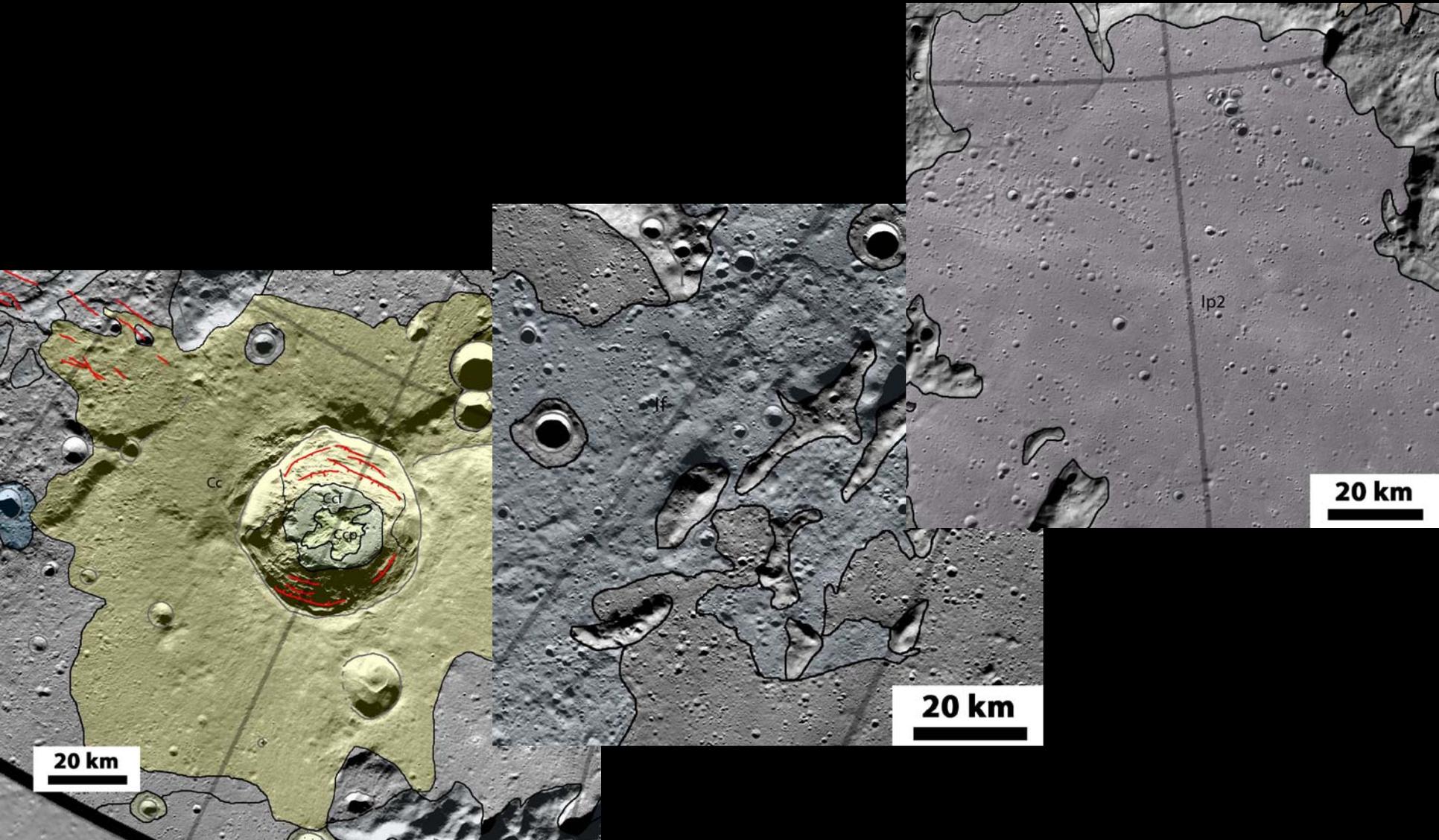
20°E

30°E

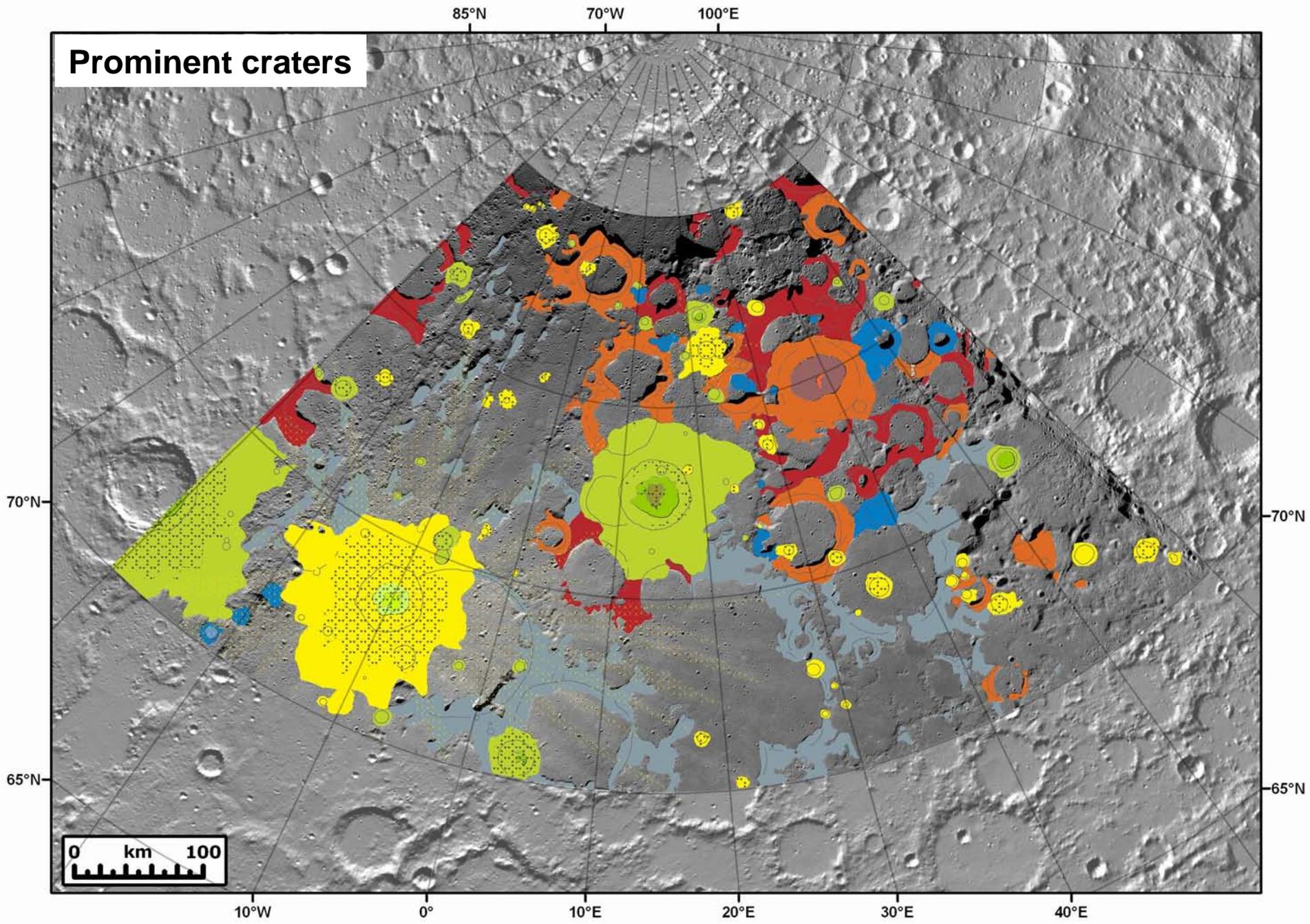
40°E

Geology (Lucchitta, 1978; I-1162)

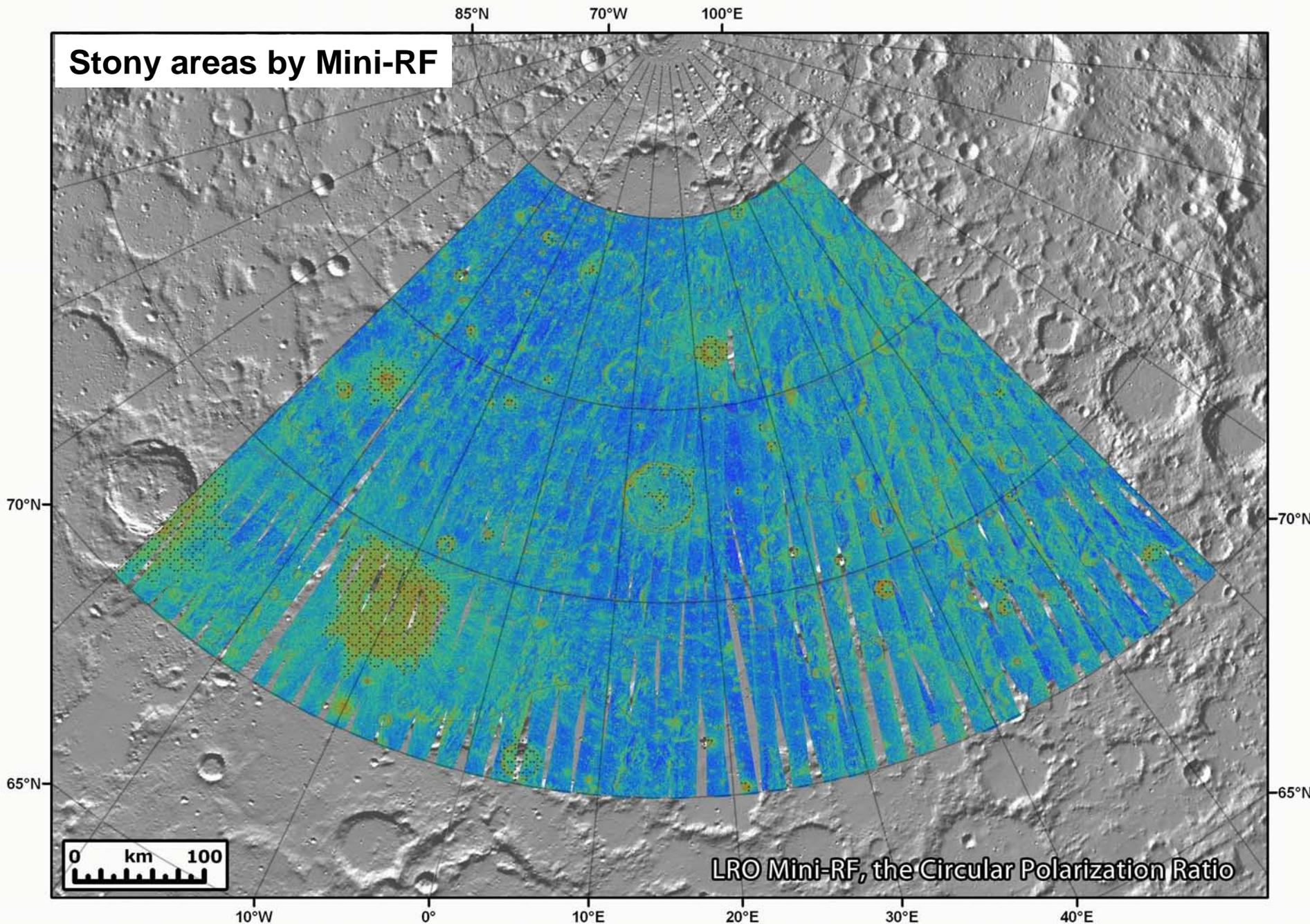
# Main Terrain materials: prominent craters, basin ejecta, plains



# Prominent craters

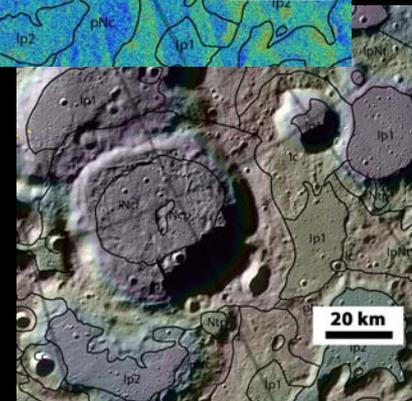
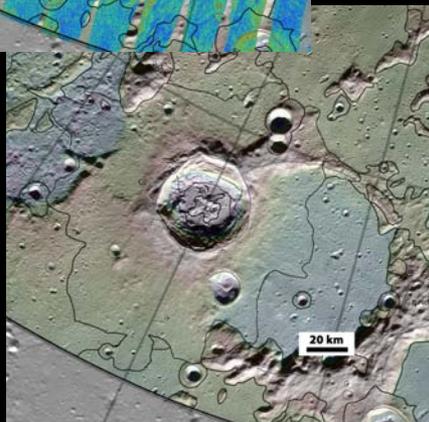
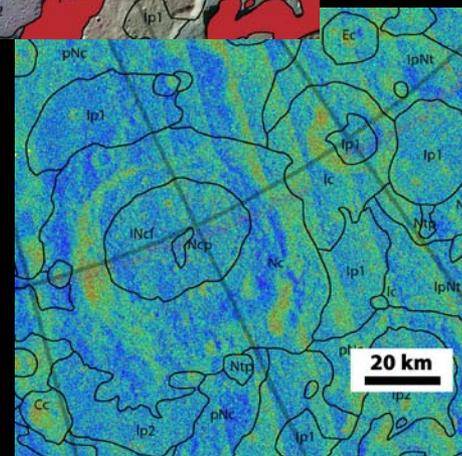
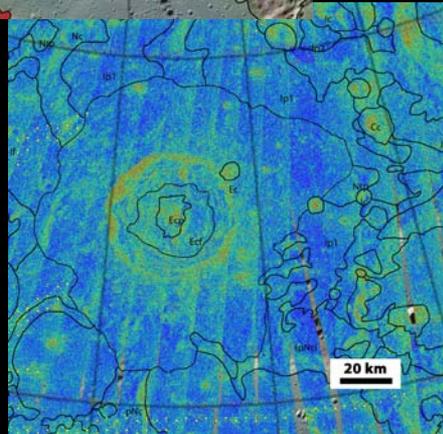
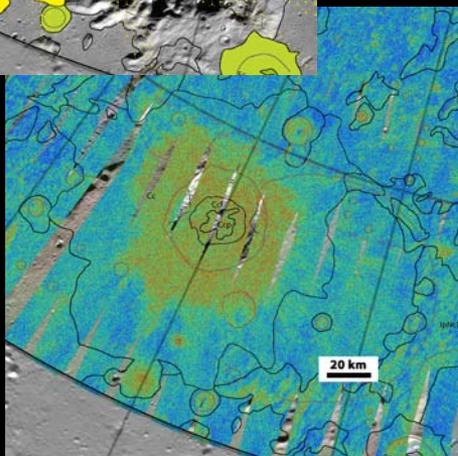
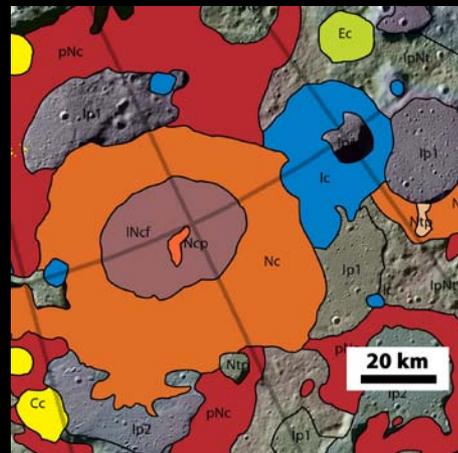
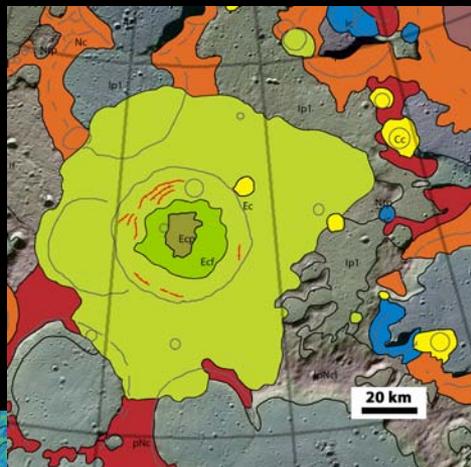
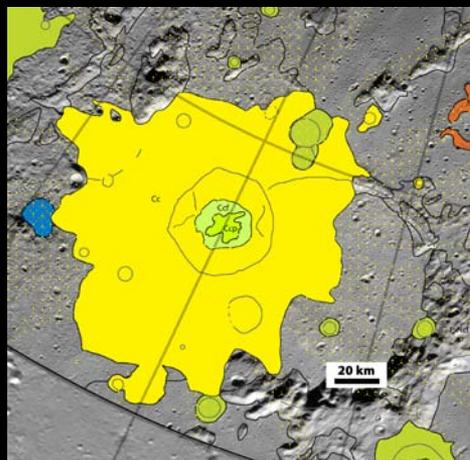


# Stony areas by Mini-RF



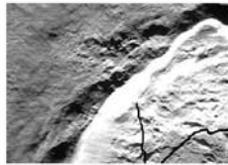
Younger ←

Older

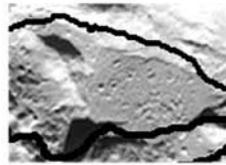


# Crater complexes from old to younger

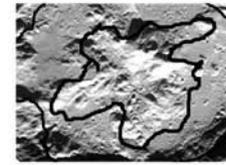
Cc



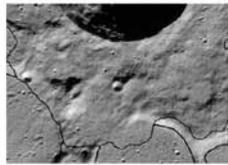
Ccf



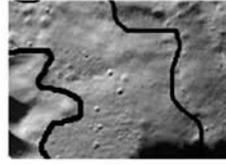
Ccp



Ec



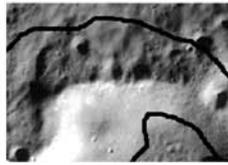
Ecf



Ecp



Ic



INcf



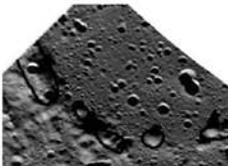
Ncp



Nc

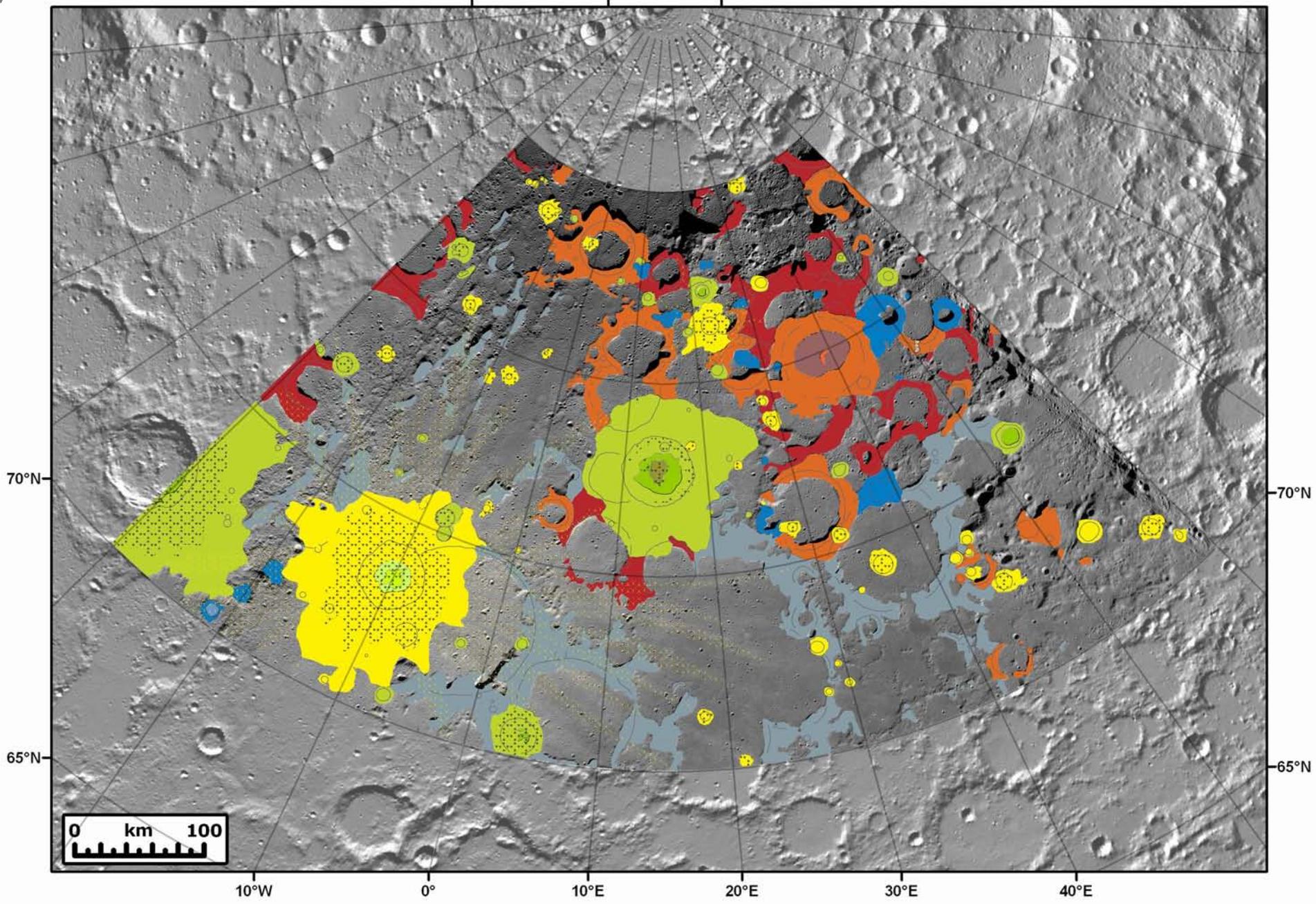


pNc





85°N 70°W 100°E



0 km 100

10°W 0° 10°E 20°E 30°E 40°E

70°N

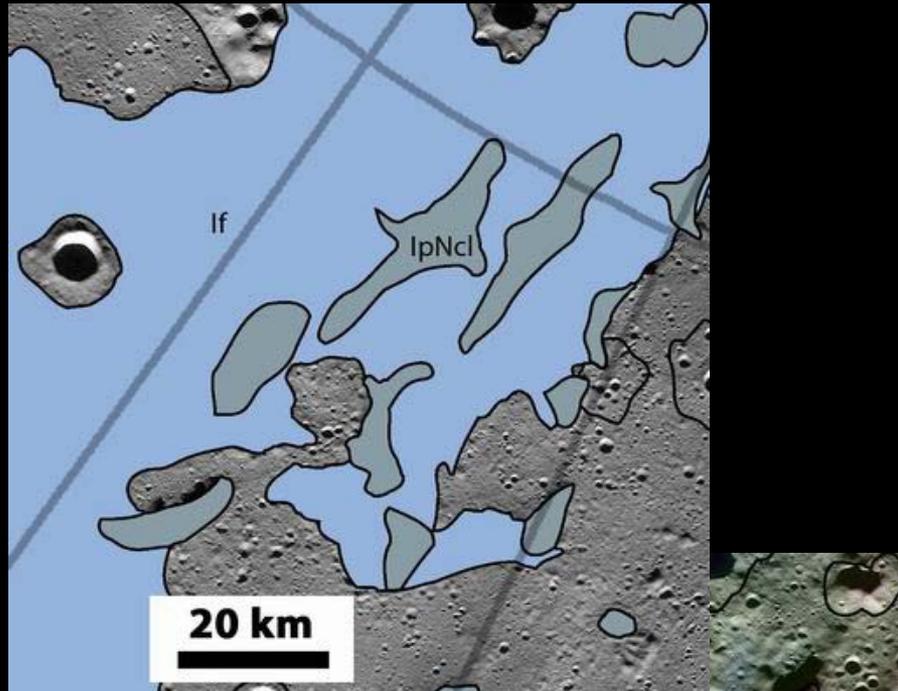
70°N

65°N

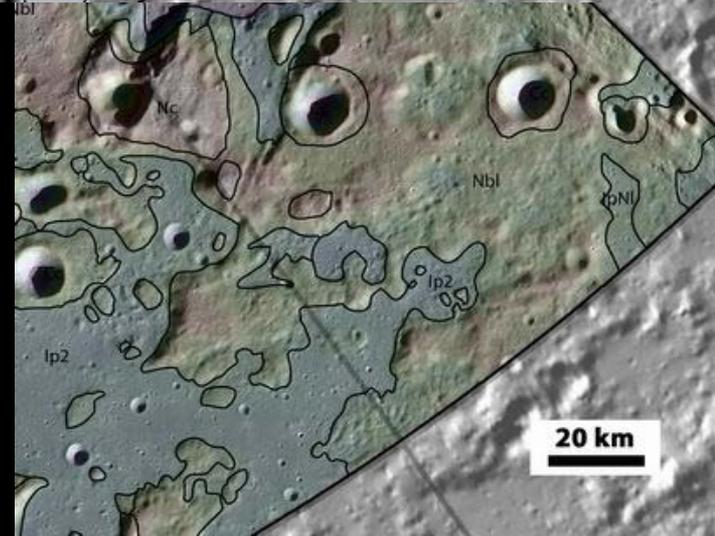
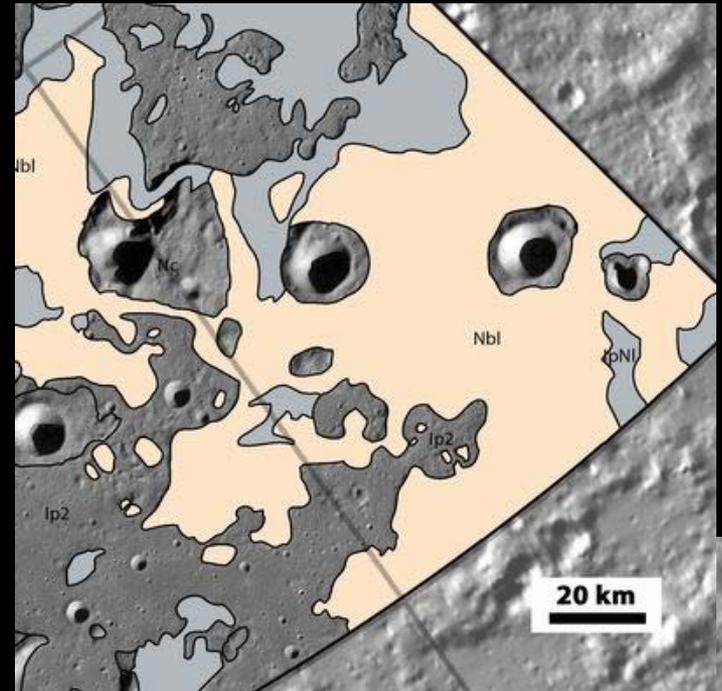
65°N

# Basin materials

## Formation Fra-Mauro If (Imbrium basin material)

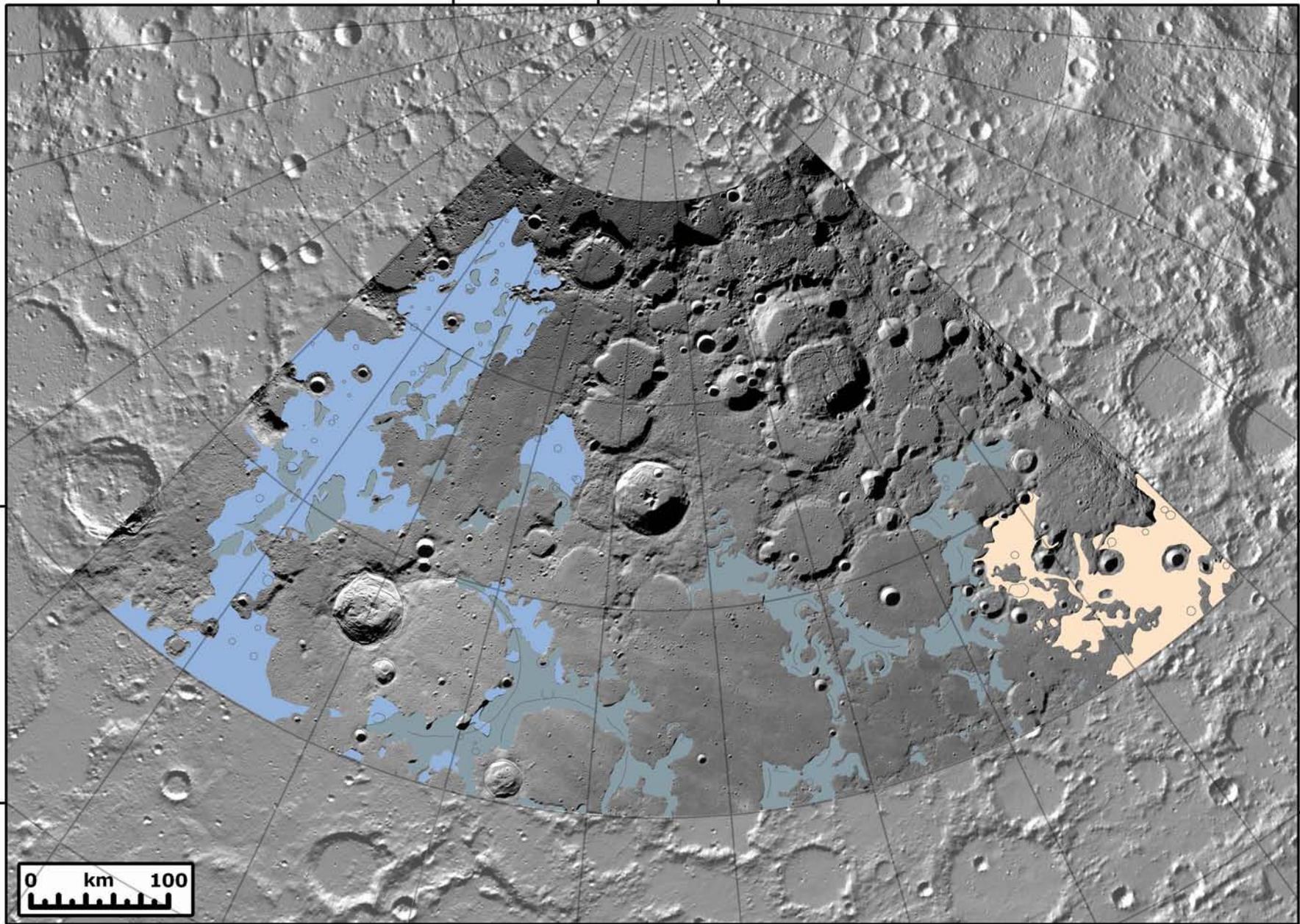


## Complex Nbl (Nectarian Humboldtianum basin material)





85°N 70°W 100°E



70°N

70°N

65°N

65°N



10°W

0°

10°E

20°E

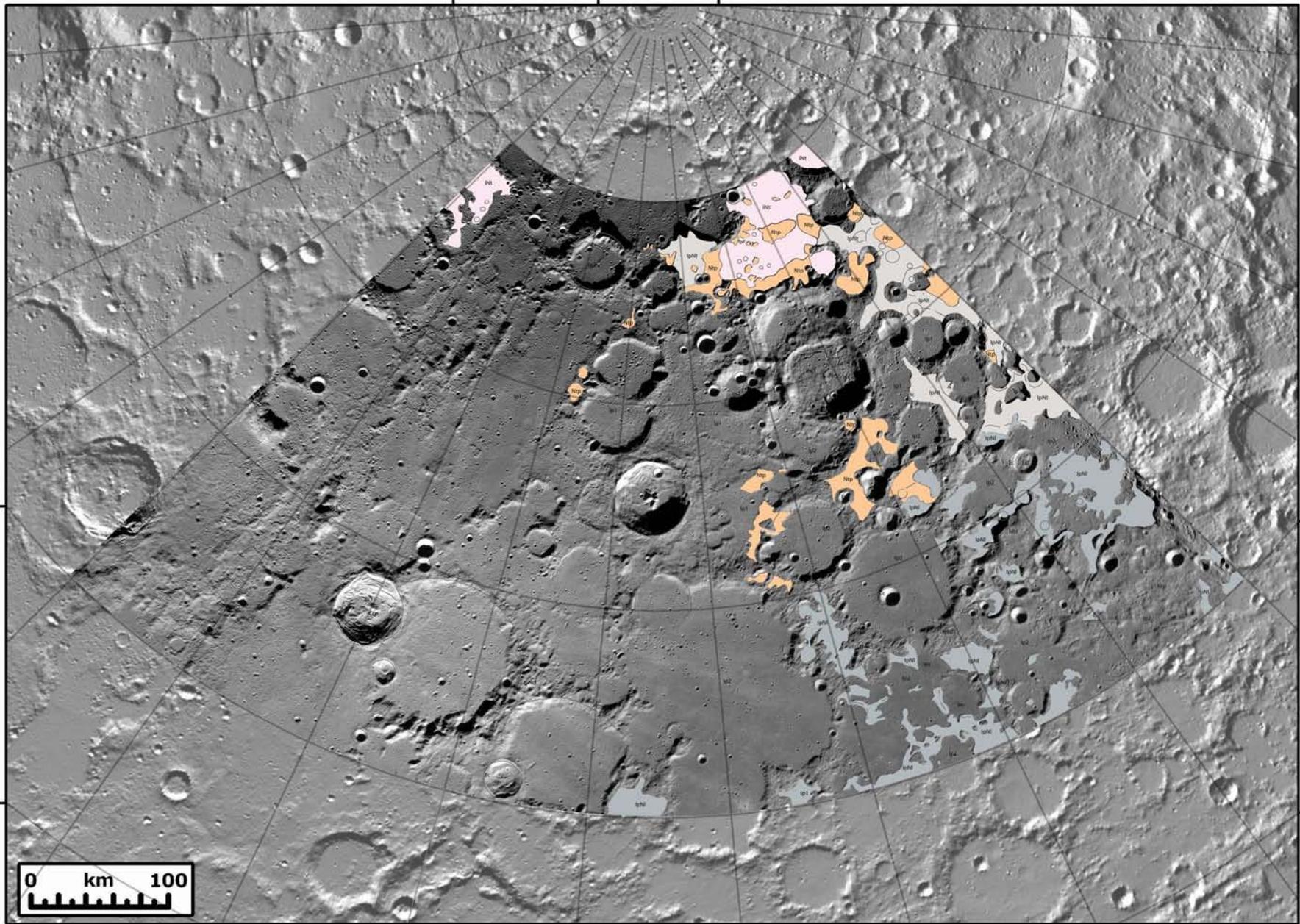
30°E

40°E





85°N 70°W 100°E



70°N

70°N

65°N

65°N



10°W

0°

10°E

20°E

30°E

40°E

# Younger highland plain materials Ip and mare material Im

Mare material Im

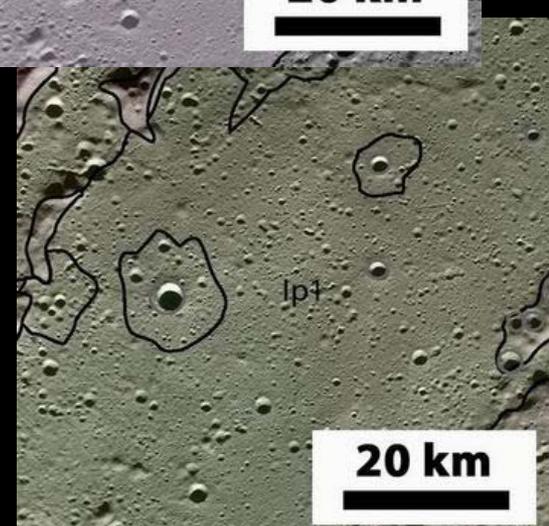
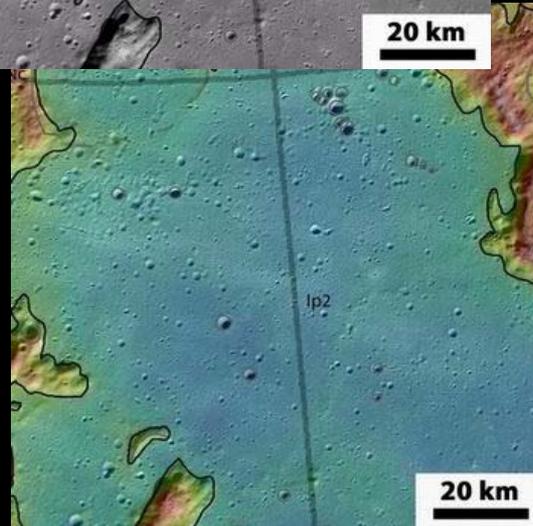
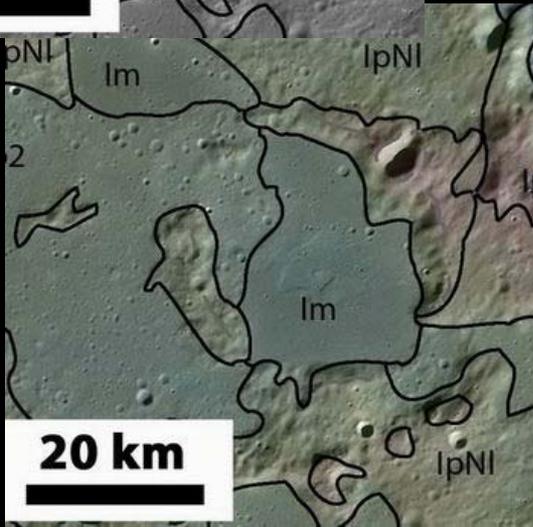
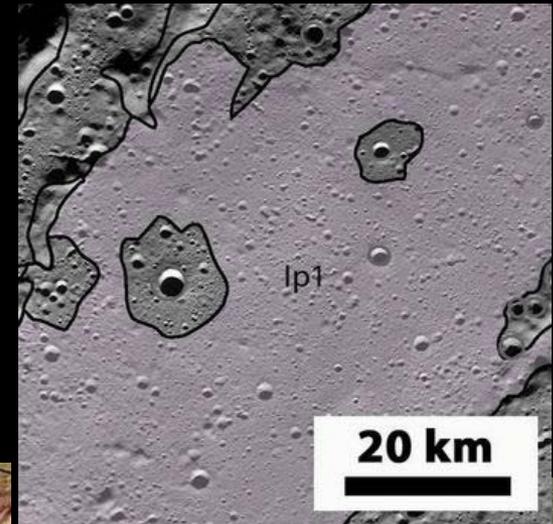
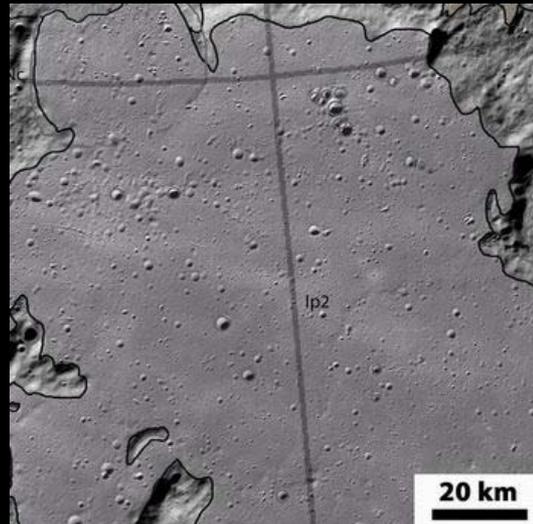
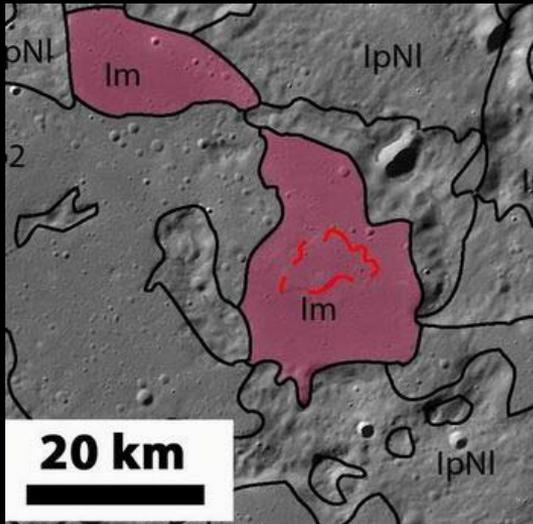
Younger highland plain material Ip

Ip2

Ip1

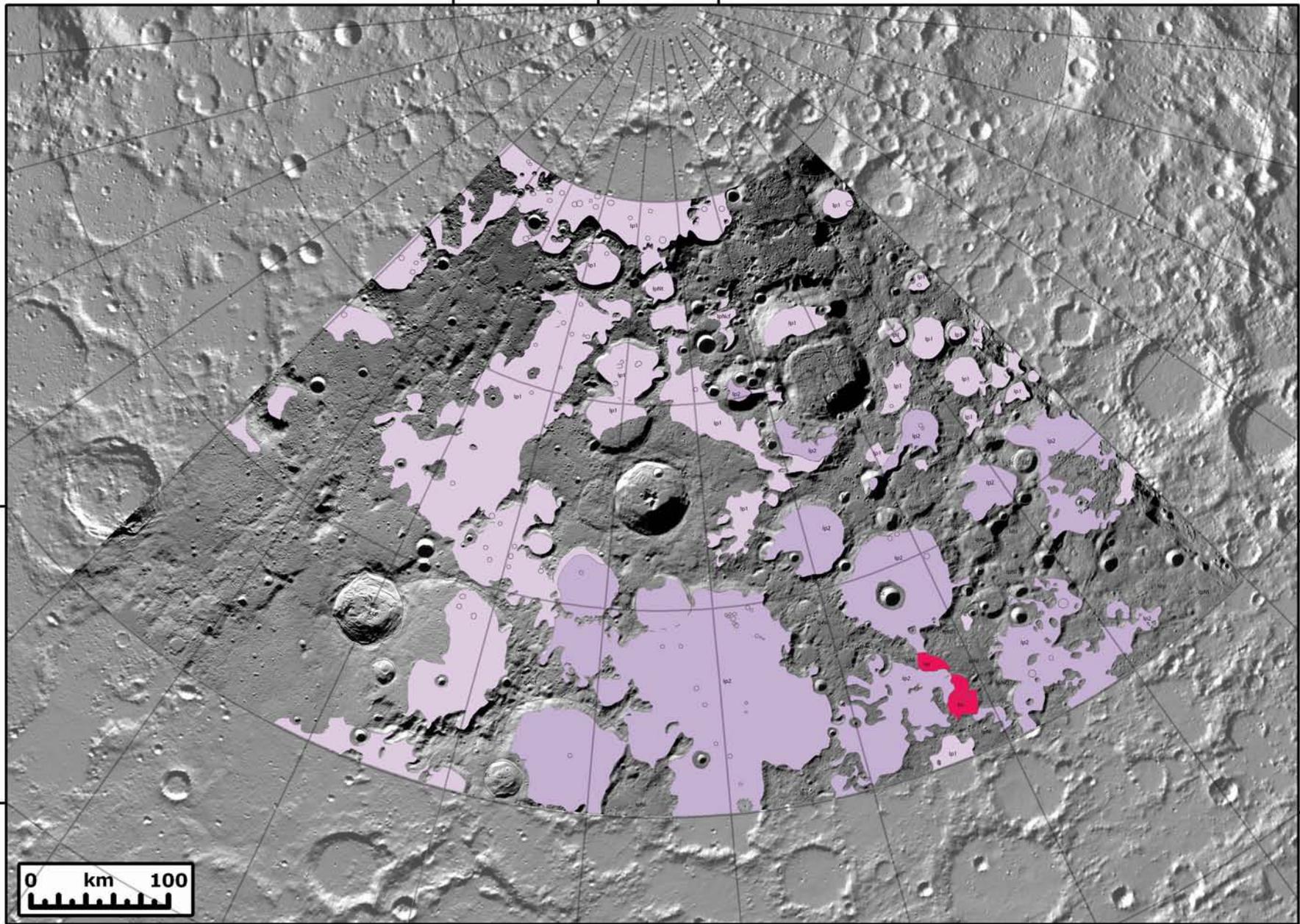
Older

Younger ←





85°N 70°W 100°E



70°N

70°N

65°N

65°N



10°W

0°

10°E

20°E

30°E

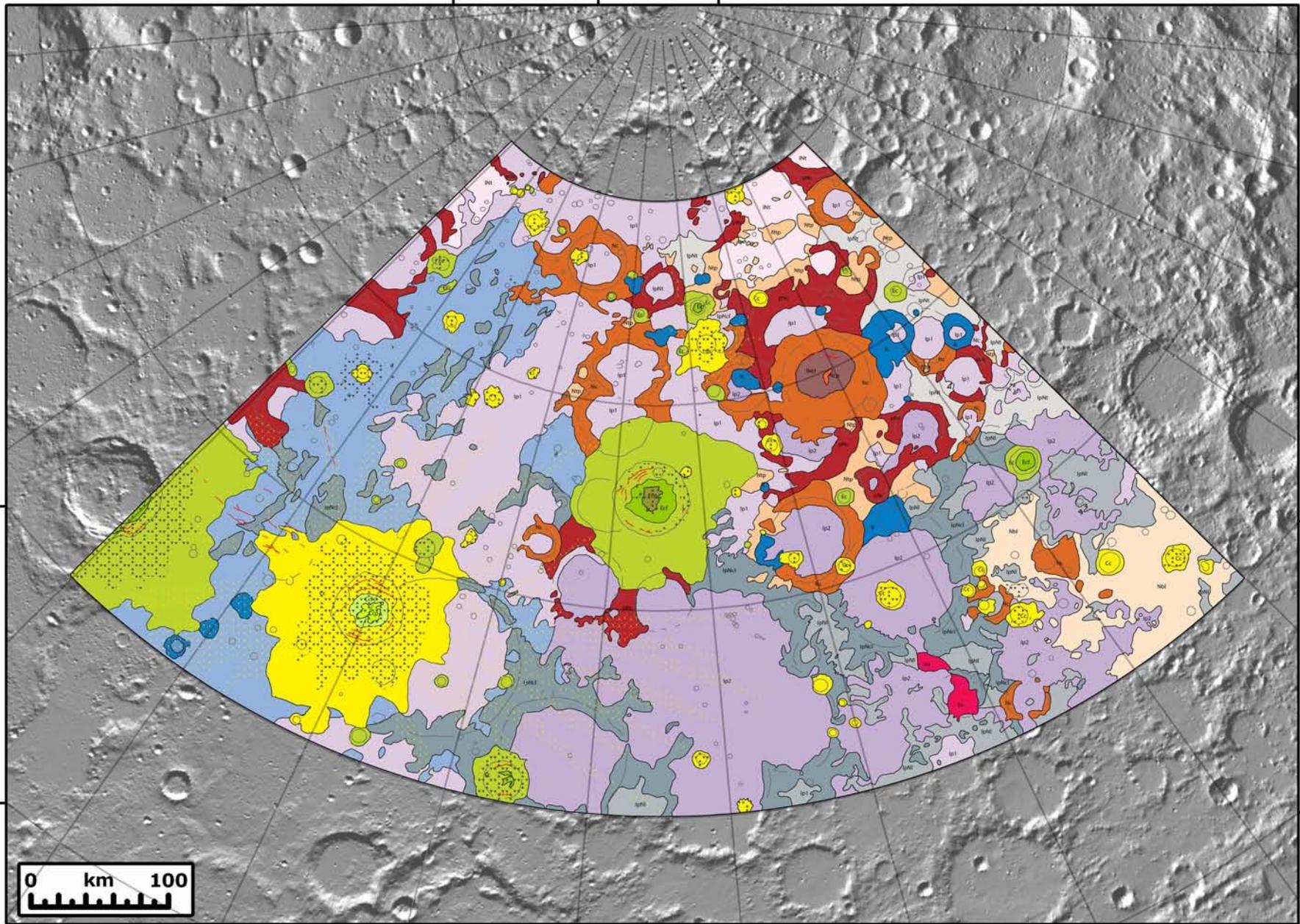
40°E



85°N

70°W

100°E



70°N

70°N

65°N

65°N



10°W

0°

10°E

20°E

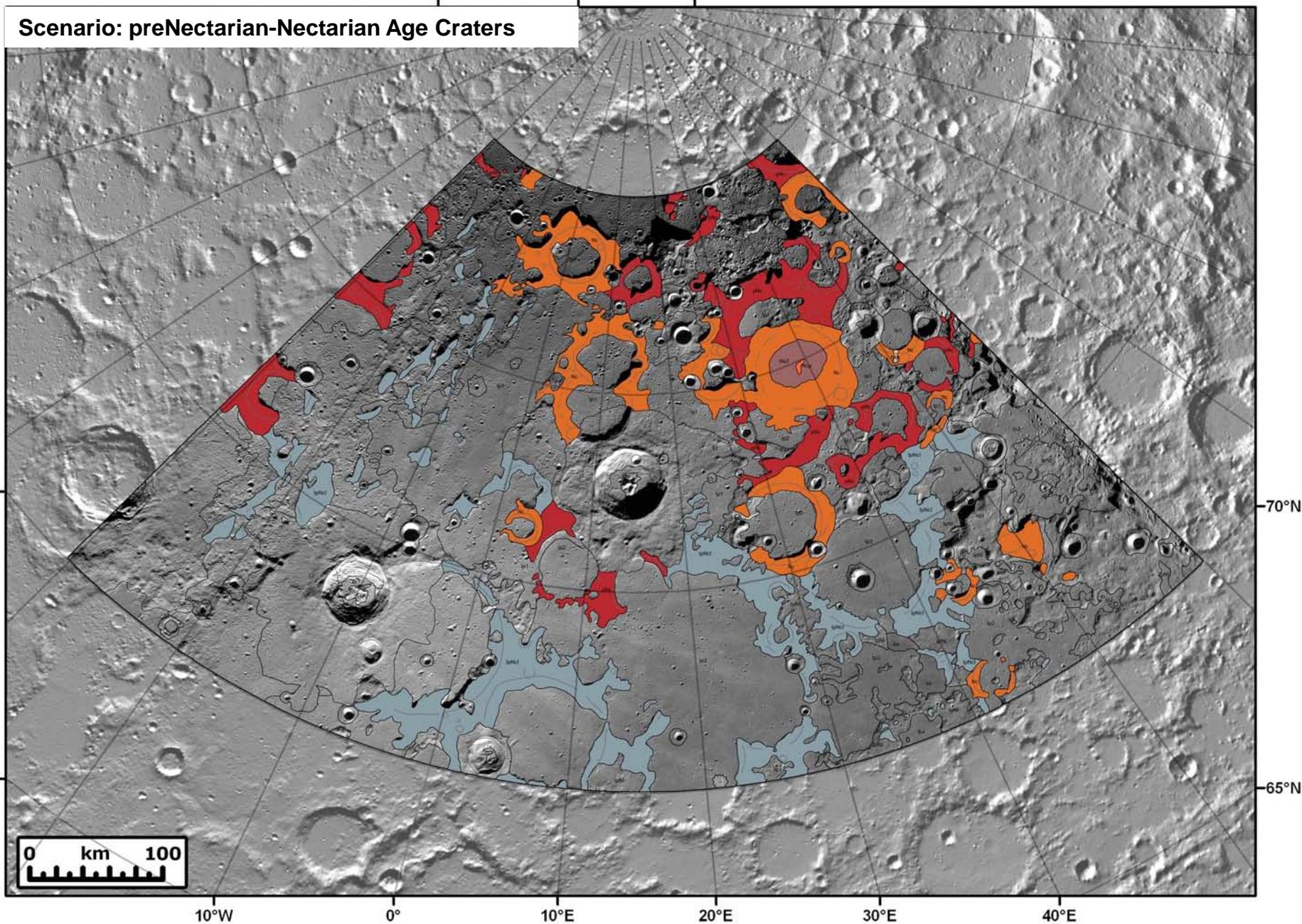
30°E

40°E



# Scenario of Geologic evolution of the studied area

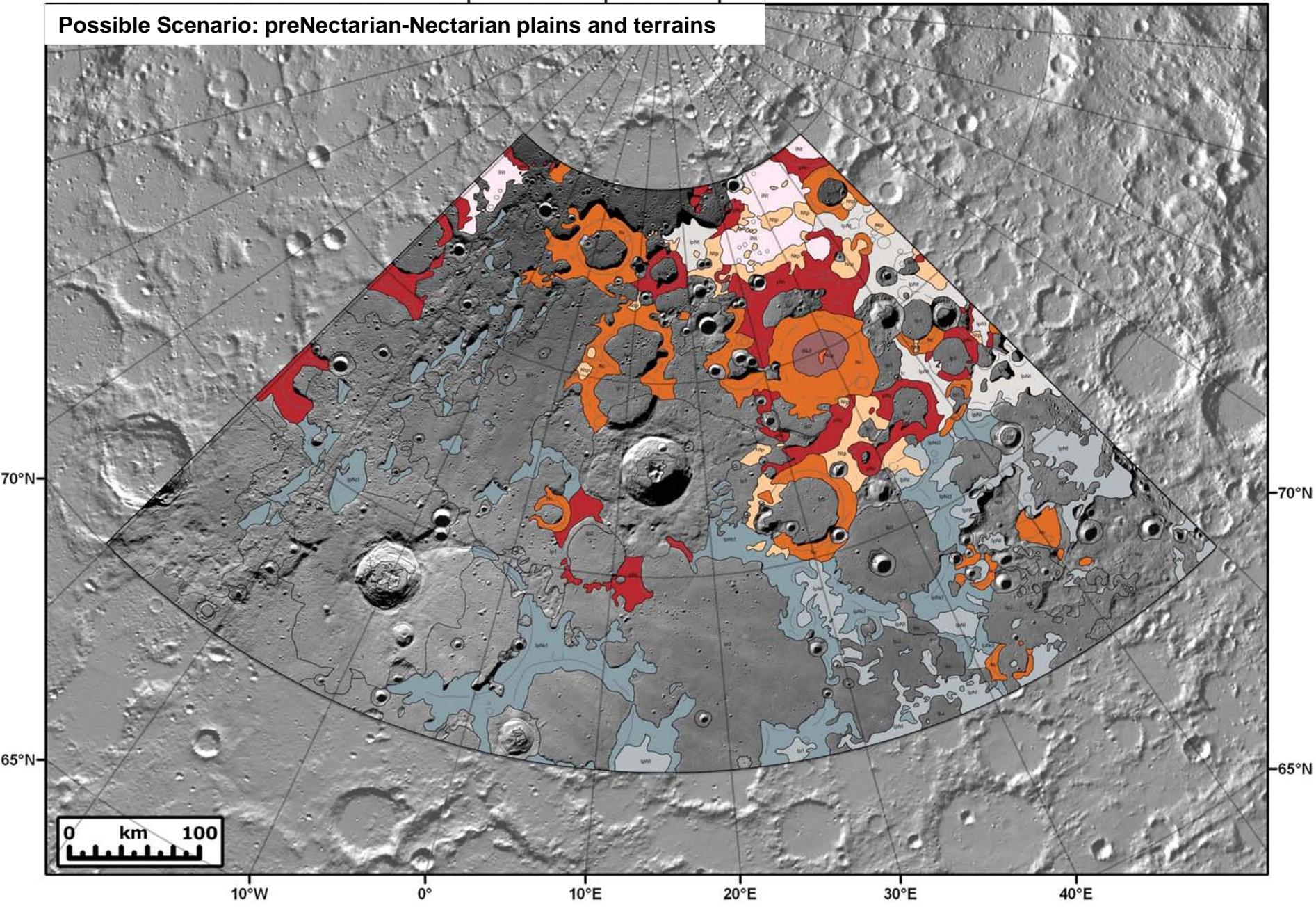
Scenario: preNectarian-Nectarian Age Craters





85°N 70°W 100°E

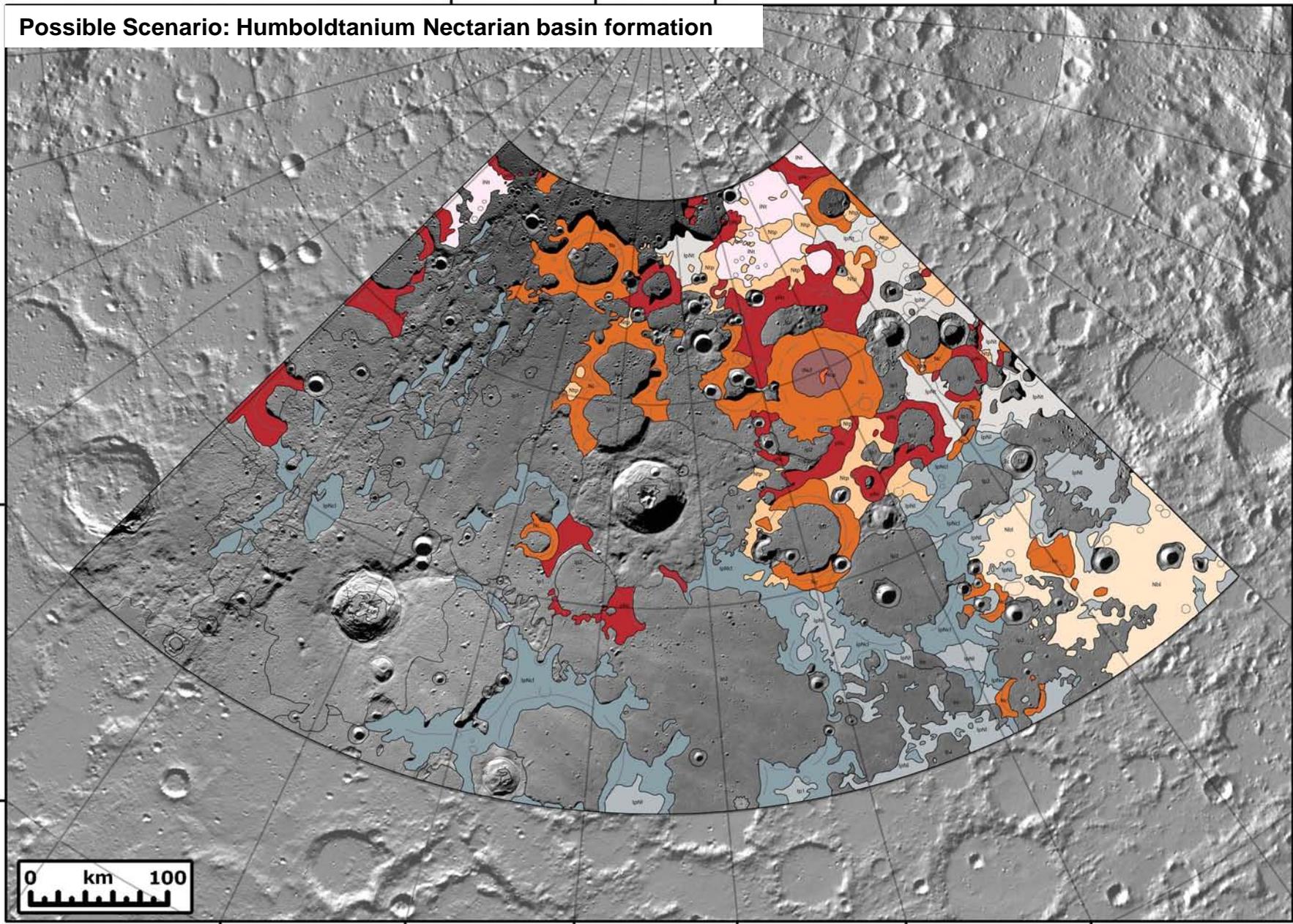
### Possible Scenario: preNectarian-Nectarian plains and terrains





85°N 70°W 100°E

### Possible Scenario: Humboldtium Nectarian basin formation



0 km 100

10°W 0° 10°E 20°E 30°E 40°E

70°N

70°N

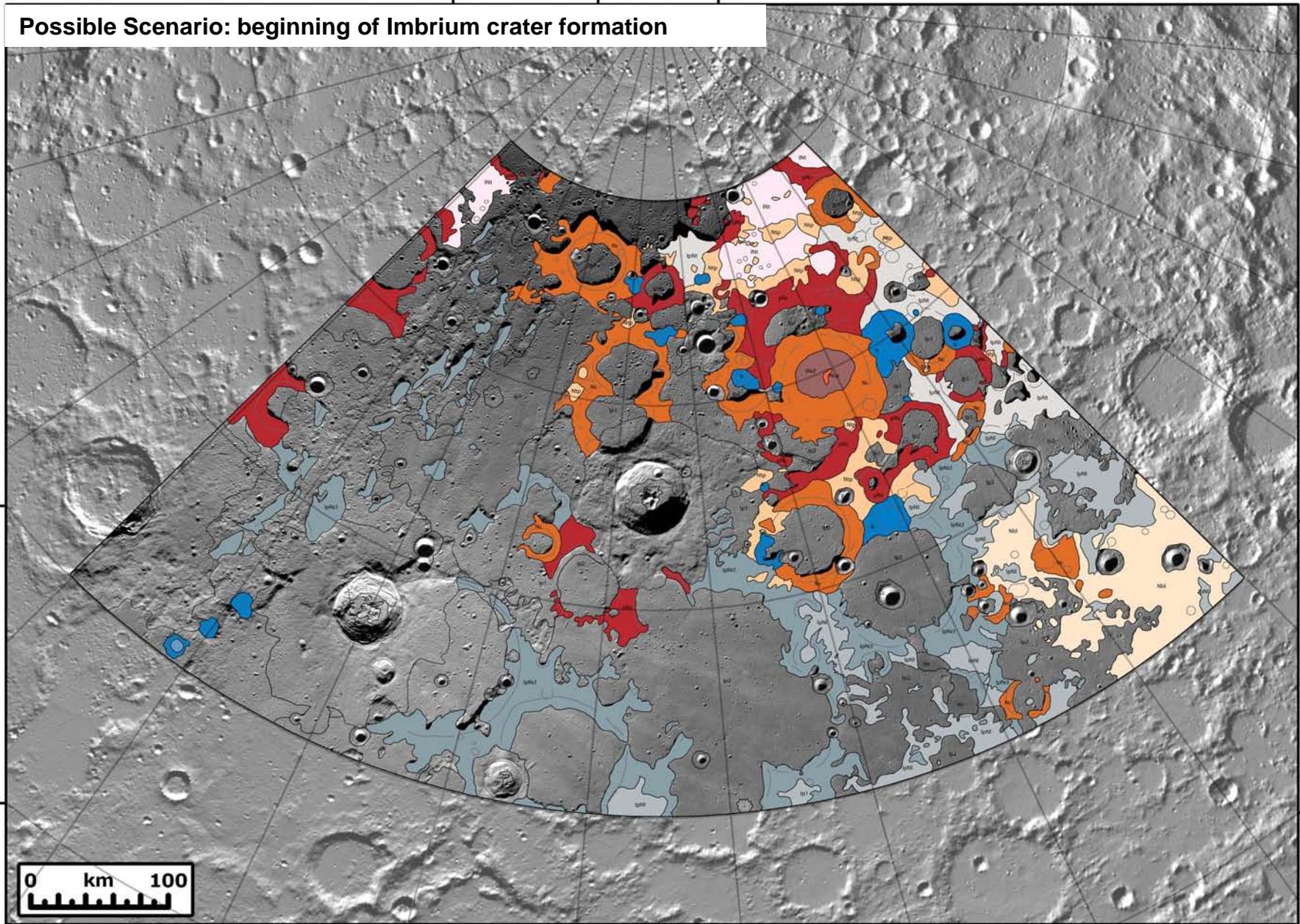
65°N

65°N



85°N 70°W 100°E

### Possible Scenario: beginning of Imbrium crater formation



10°W 0° 10°E 20°E 30°E 40°E

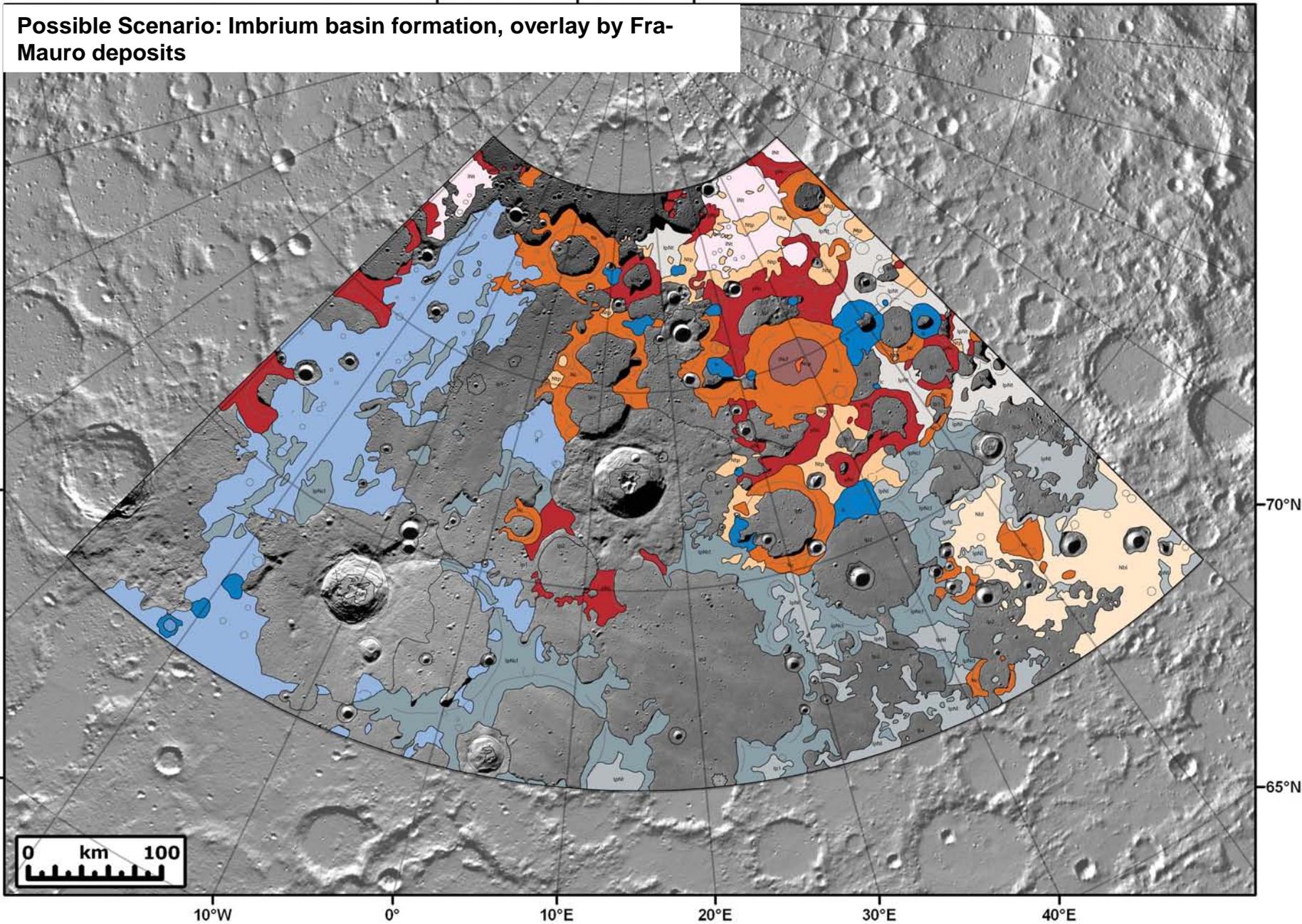
70°N

70°N

65°N

65°N

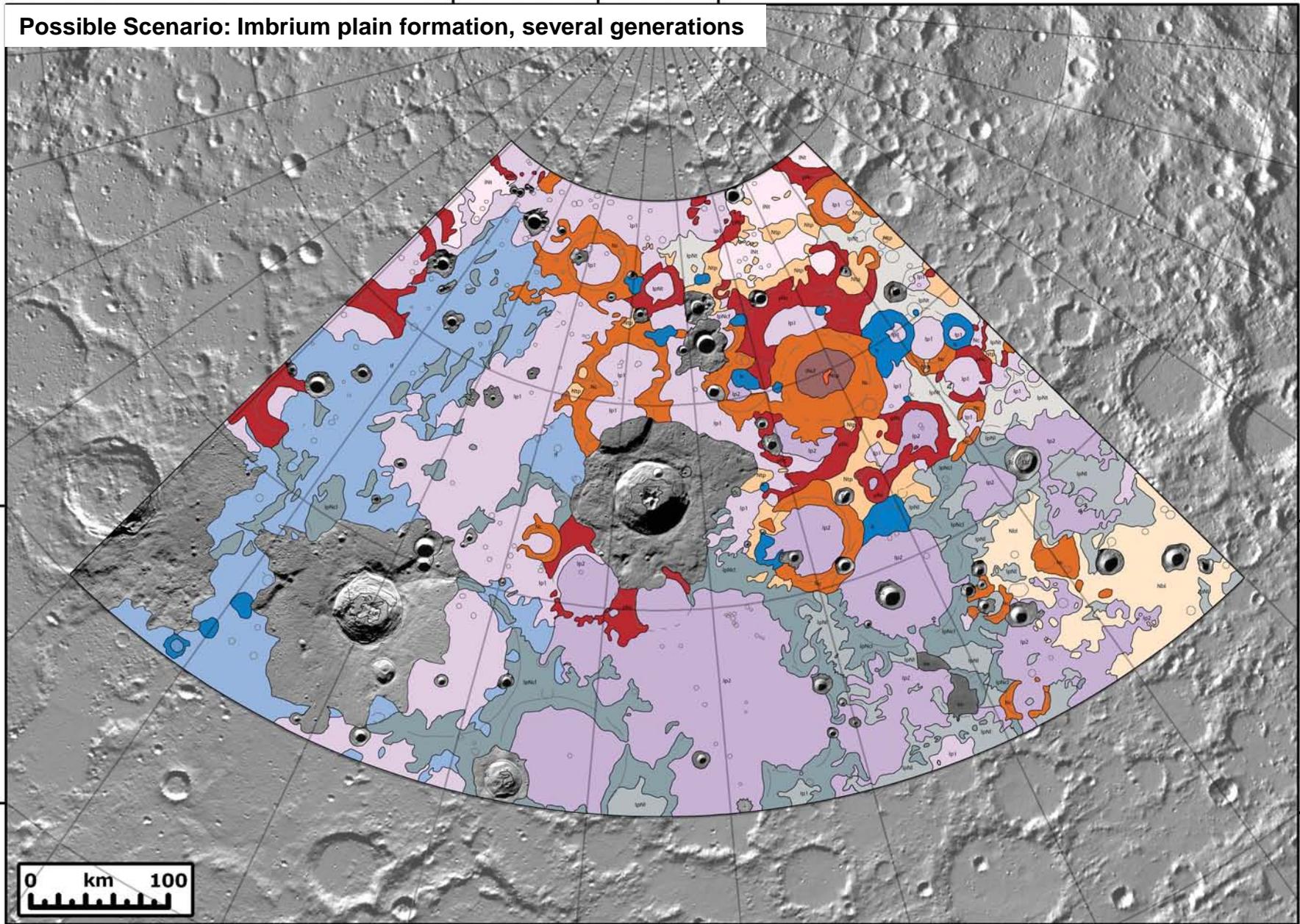
Possible Scenario: Imbrium basin formation, overlay by Fra-Mauro deposits





85°N 70°W 100°E

### Possible Scenario: Imbrium plain formation, several generations



0 km 100

10°W 0° 10°E 20°E 30°E 40°E

70°N

70°N

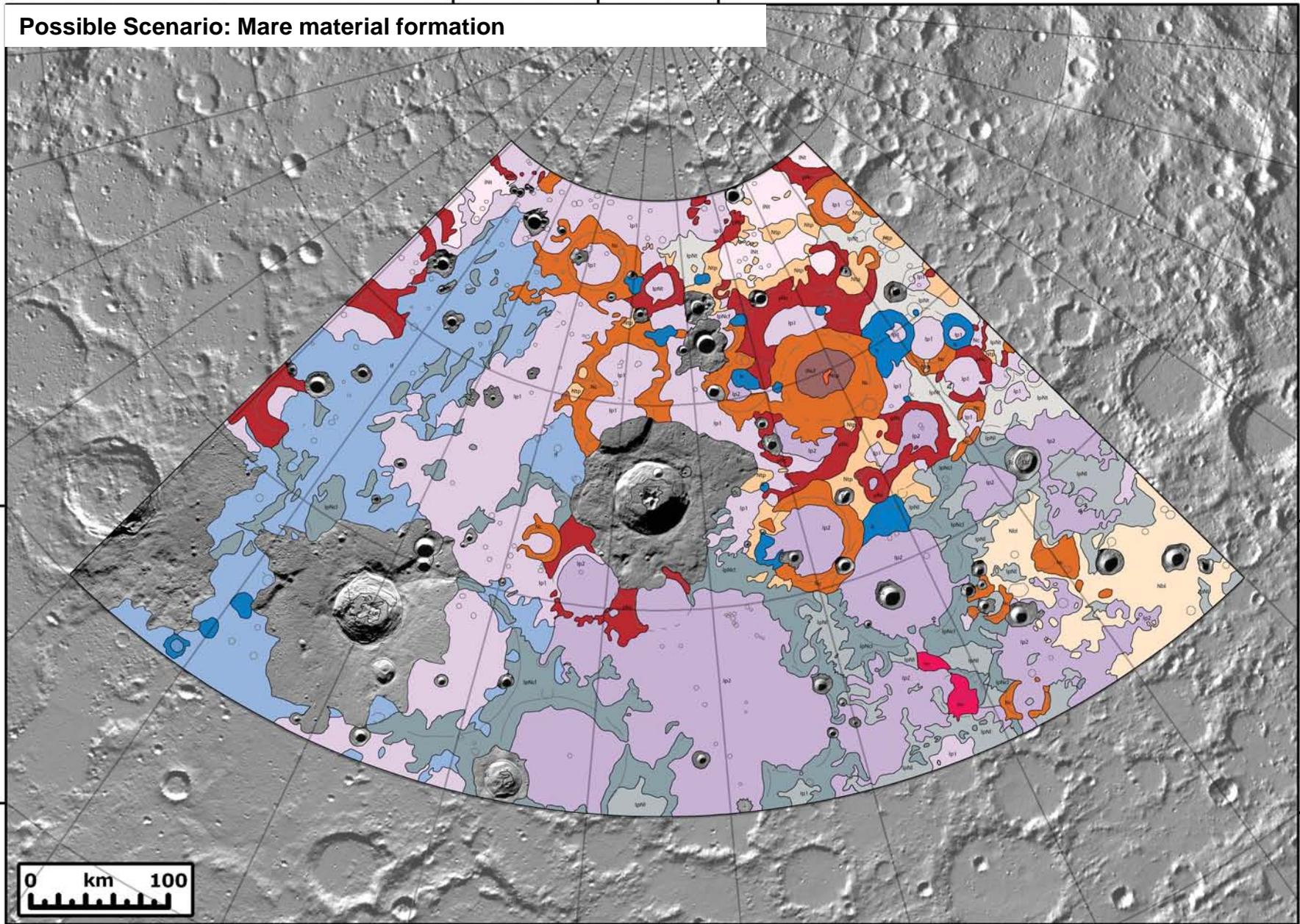
65°N

65°N



85°N 70°W 100°E

### Possible Scenario: Mare material formation



70°N

70°N

65°N

65°N



10°W

0°

10°E

20°E

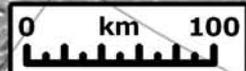
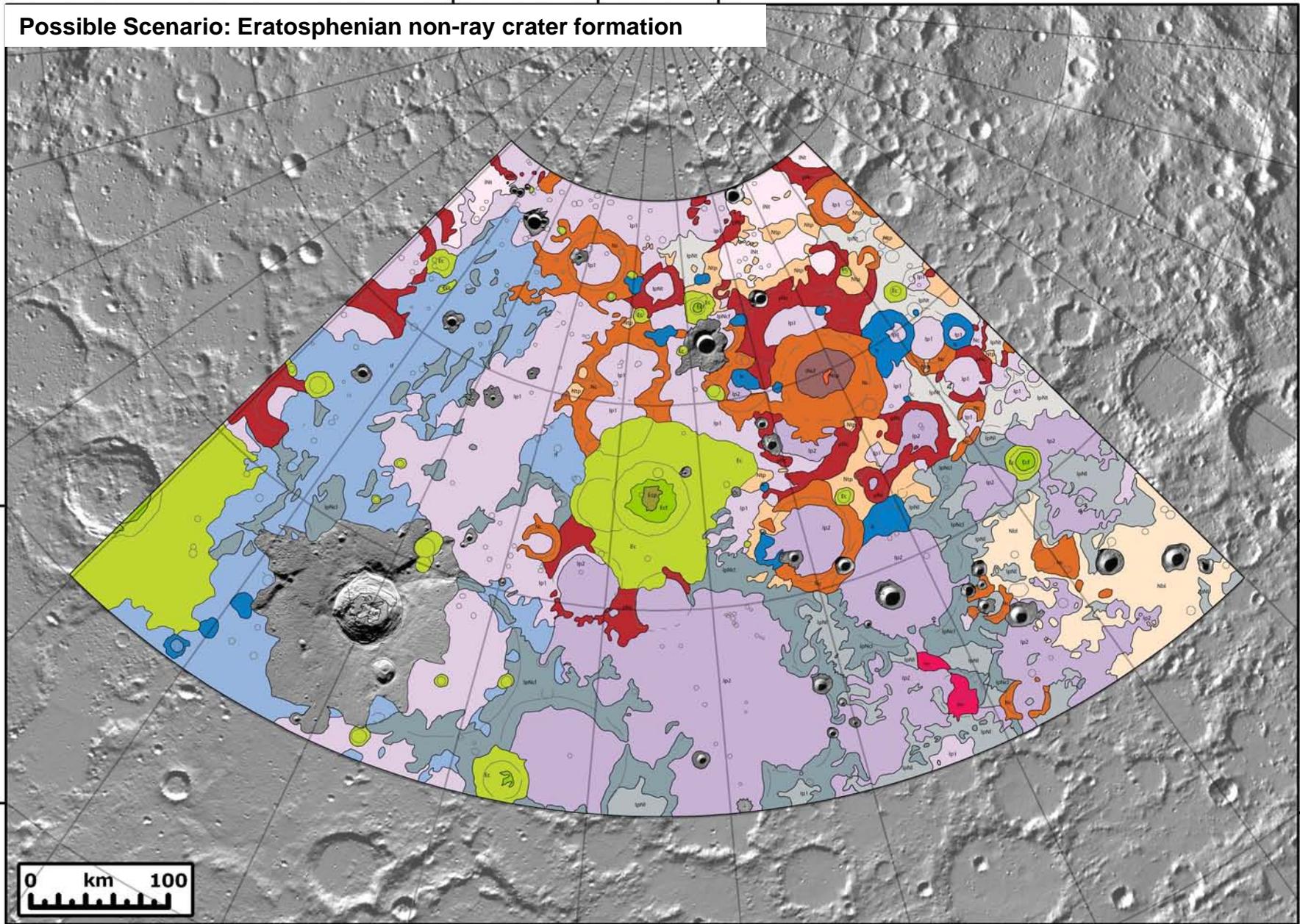
30°E

40°E



85°N 70°W 100°E

# Possible Scenario: Eratosphenian non-ray crater formation



10°W 0° 10°E 20°E 30°E 40°E

70°N

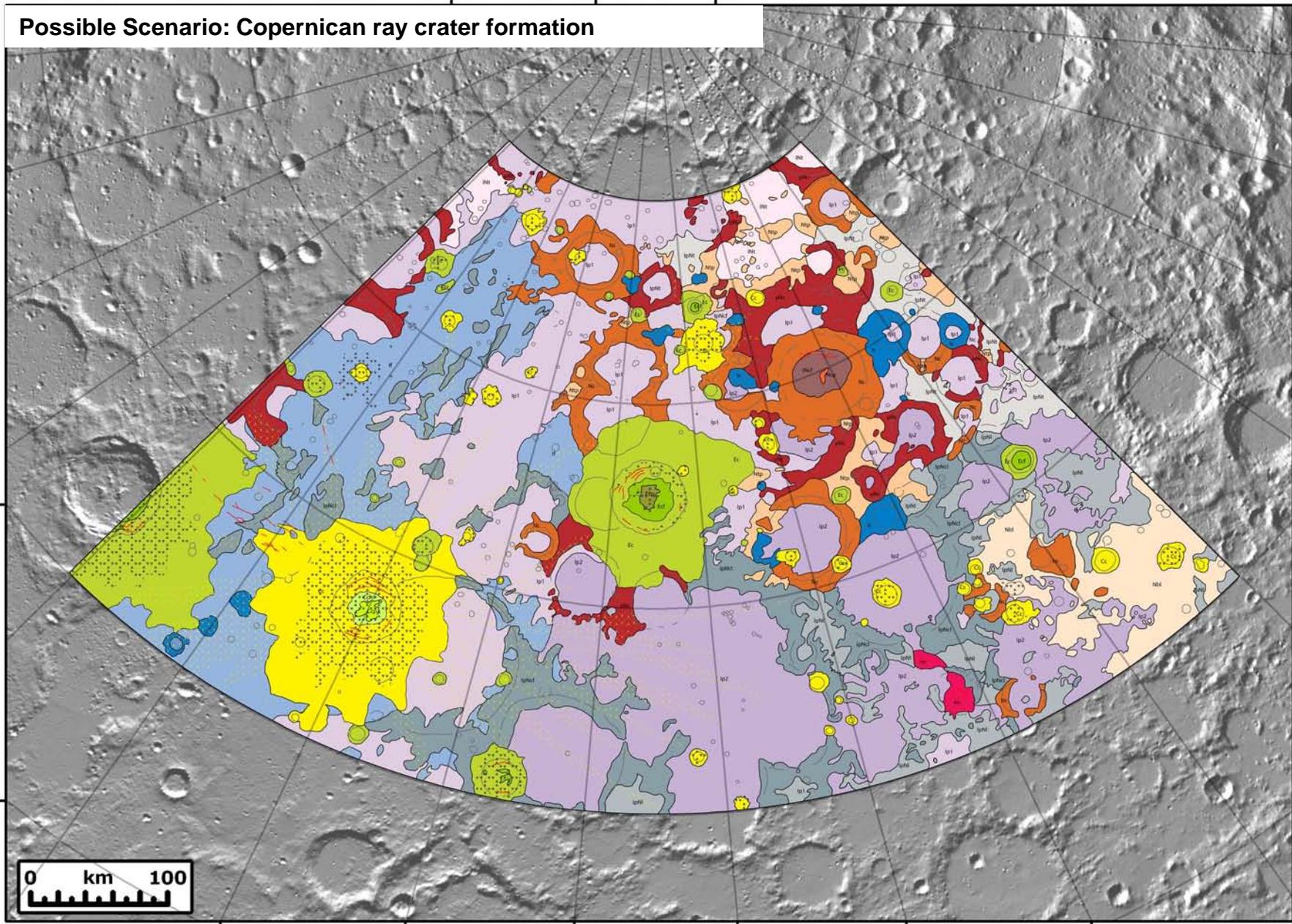
70°N

65°N

65°N

Possible Scenario: Copernican ray crater formation

85°N 70°W 100°E



70°N

70°N

65°N

65°N



10°W

0°

10°E

20°E

30°E

40°E

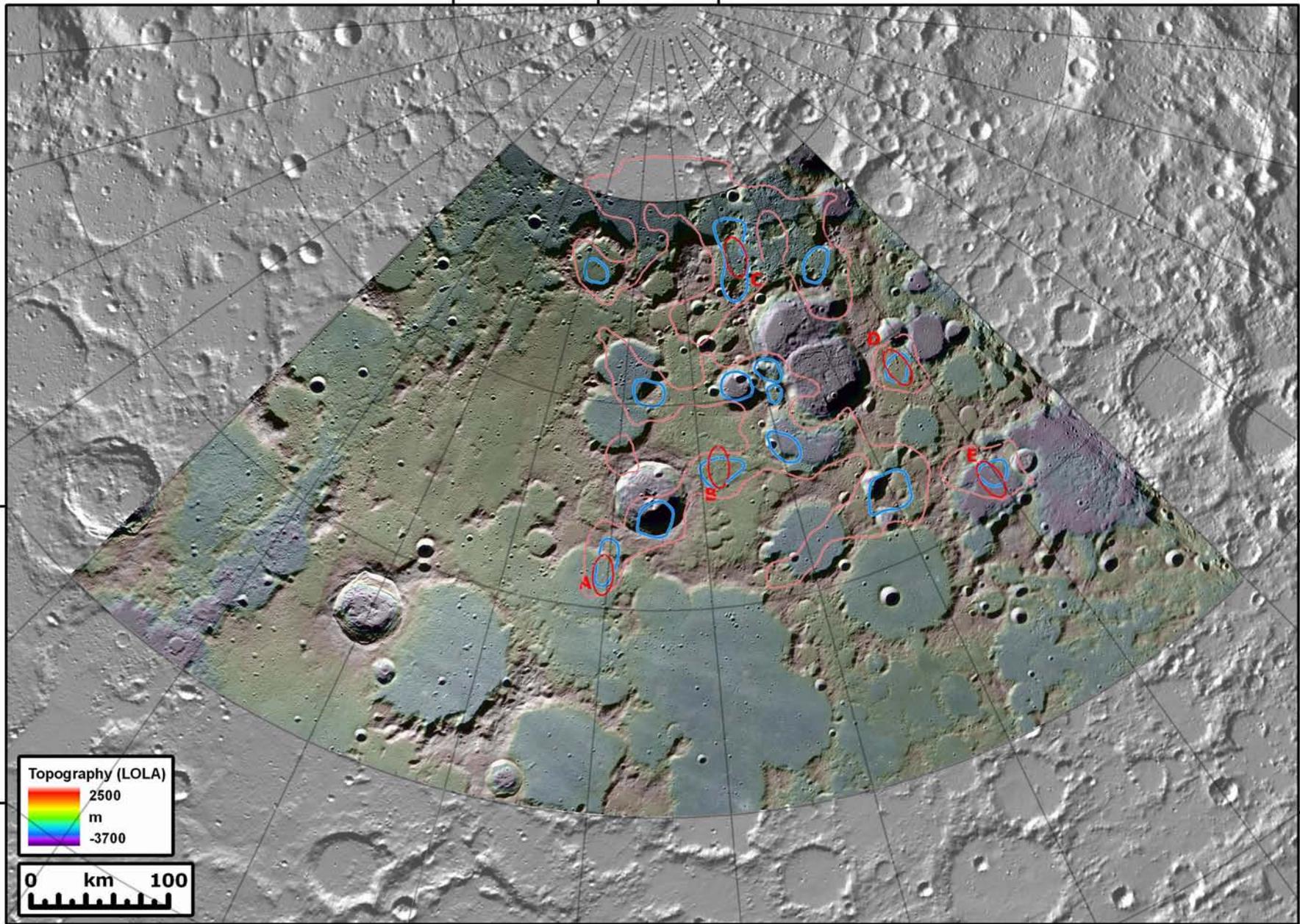


# Criteria of the Luna Glob Candidate Landing Site selection:

1. NPOL Ballistic constraints: The boundaries of landing area:  
 $70-85^{\circ}$  N,  $-30 - 60^{\circ}$  E
2. High Volatile Enrichment (by LEND data, Low neutron flux).
3. Safety of approaching trace (LOLA data).
4. The landing area (15x30 km ellipse) should be safely smooth (LROC, LOLA, Micro-RF data).



85°N 70°W 100°E

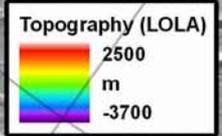


70°N

70°N

65°N

65°N



10°W 0° 10°E 20°E 30°E 40°E



85°N 70°W 100°E

70°N

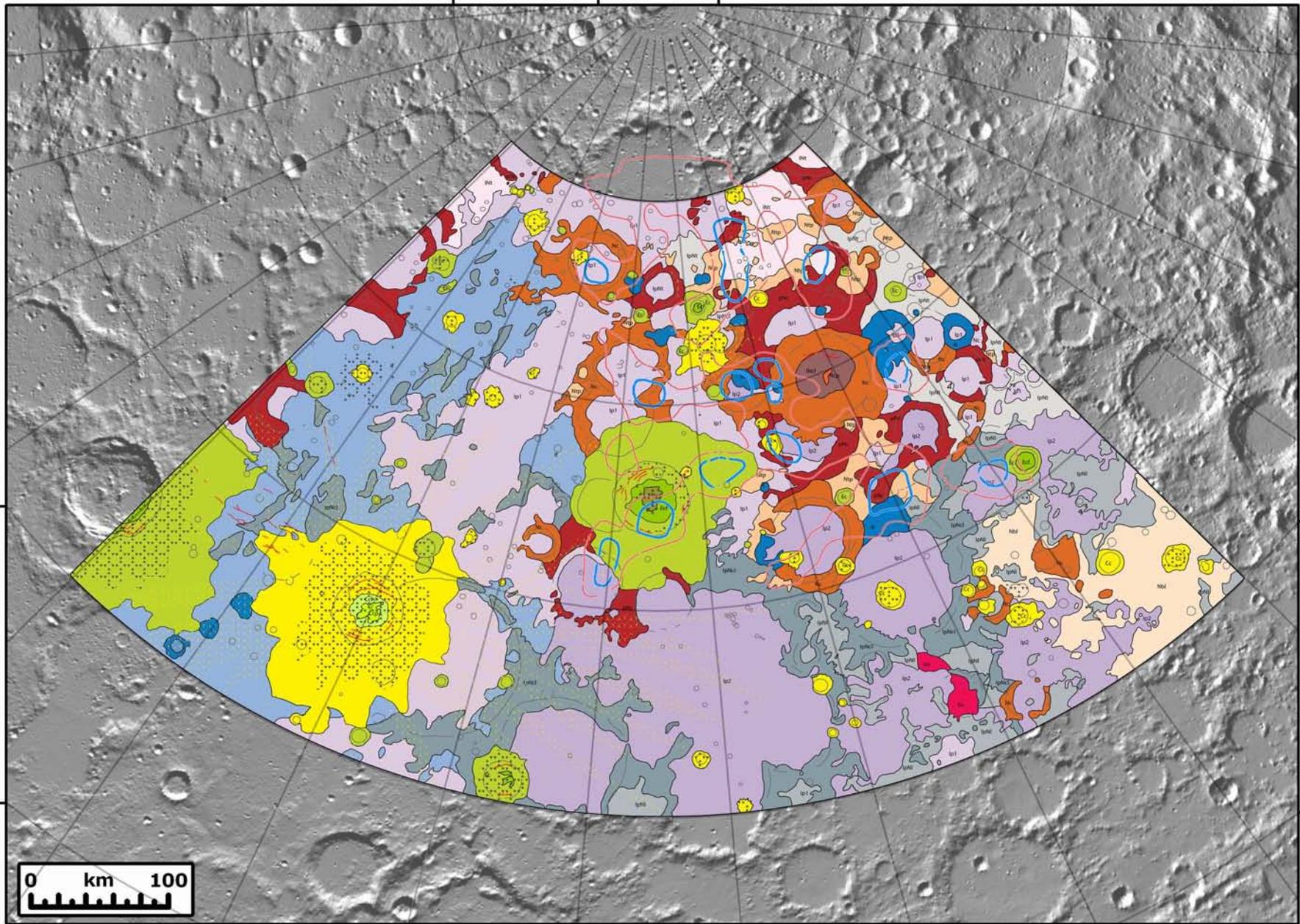
70°N

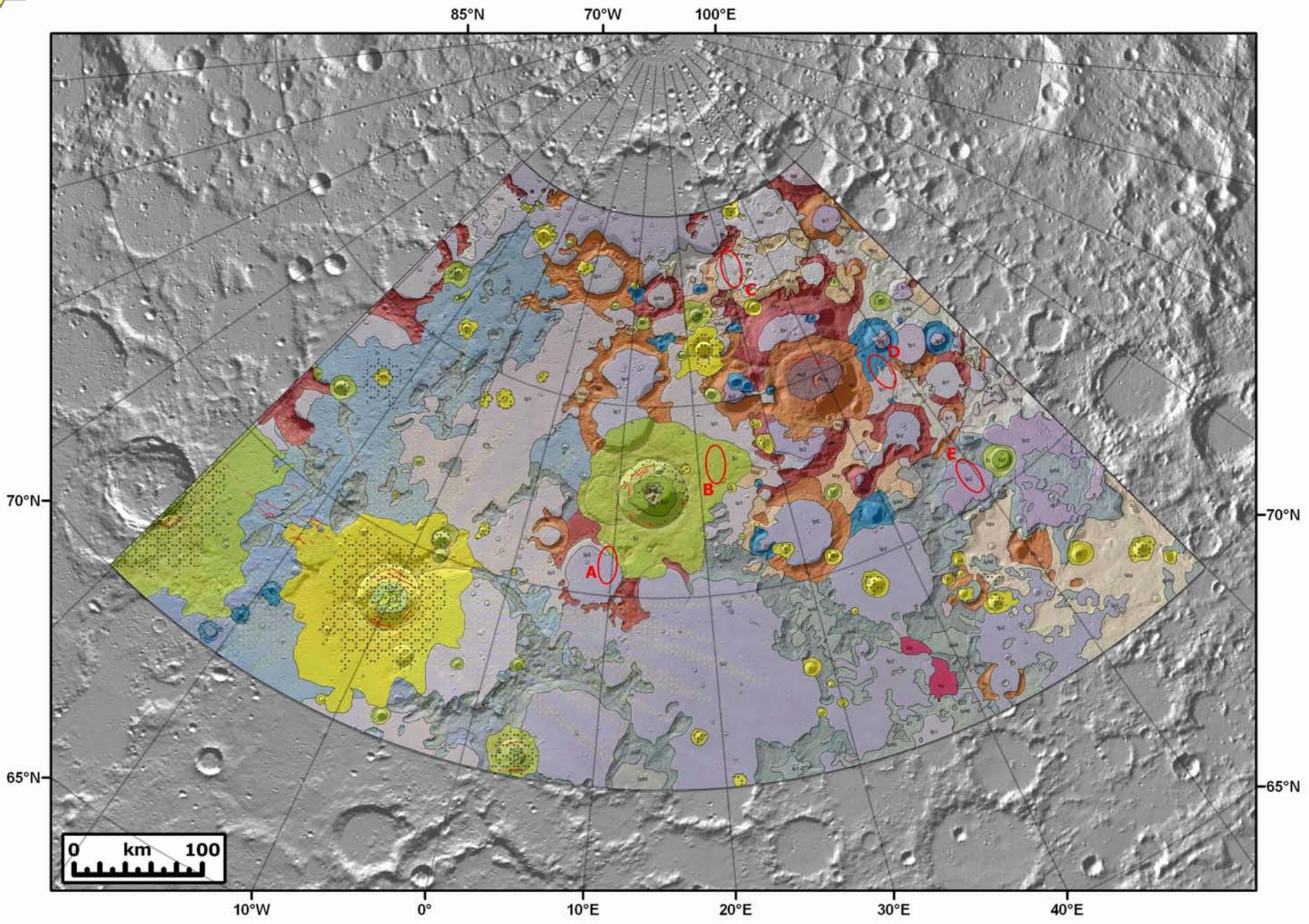
65°N

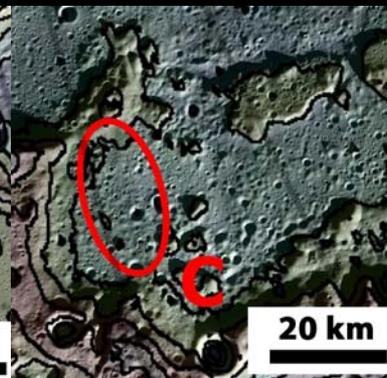
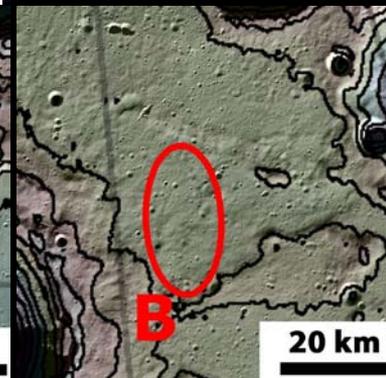
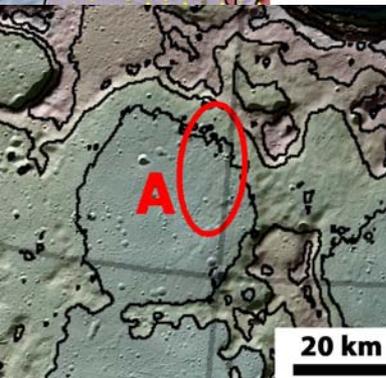
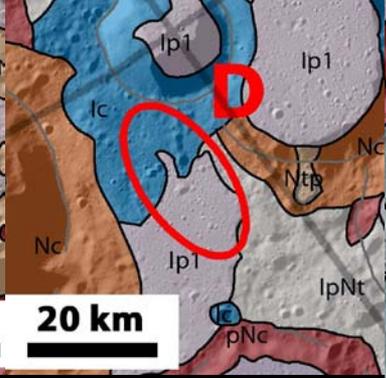
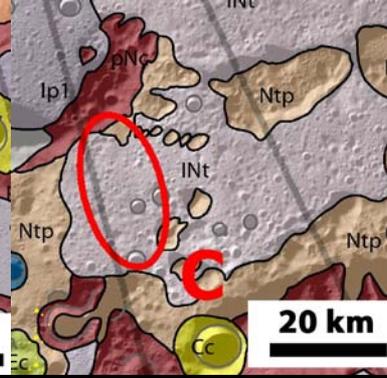
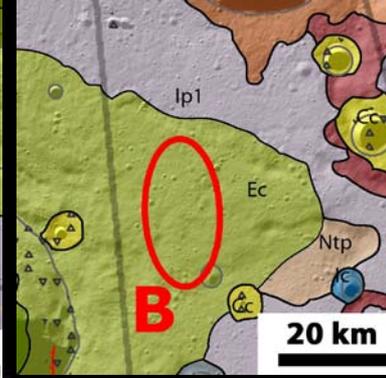
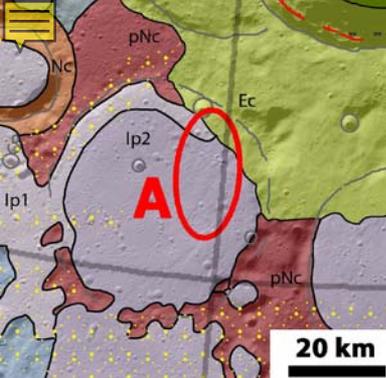
65°N



10°W 0° 10°E 20°E 30°E 40°E







<b>A north</b> <b>75.8° N, 9.5° E</b>	<b>B north</b> <b>78.4° N, 22.2° E</b>	<b>C north</b> <b>83.3° N, 31.6° E</b>	<b>D north</b> <b>79.3° N, 47.7° E</b>	<b>E north</b> <b>75.7° N, 49.3° E</b>
Plains Ip2, (Rays of Cc ?)	Crater material Ec (possible Ip1 ?)	Material of hum- mocky terrain INt	Material of plains Ip1, craters Ic	Material of plains Ip2
The Floor of big crater filled by plain material Ip2	Outflow of craters	Oldest smoothed terrain	Old Intracrater terrain area filled by plain material Ip1	The Floor of big crater filled by plain material Ip2
Relatively good approach	Good approach	<b>Danger Slopes Before landing</b>	Relatively good approach	<b>Danger N Slopes Before</b>
Highland mineralogy	Highland mineralogy	Highland mineralogy	Highland mineralogy	Highland mineralogy



# Conclusions

- There is no correlation between Hydrogen anomalies and specific geologic complexes and position
- Preferable landing point material: A-North – material of imbrium highland plains Ip1, B-North – material of outflow of moderately young crater Ec



## *Acknowledgements:*

We are doing selection of landing sites on different bodies of the Solar system since late 60's, but it is the first time when we have so good data for that. These data are being provided by our coauthors who specially process them for us and we express our gratitude to these **colleagues** and **institutions** which support their and thus our activity. They are:

- **Department of Geological Sciences, Brown University, Providence, RI, USA**
- **Goddard Space Flight Center, Greenbelt, MD, USA**
- **Earth, Atmospheric and Planetary Sciences, MIT, Cambridge, MA, USA**
- **The John Hopkins University Applied Physics Lab, Laurel, MD 20723**
- **NASA USA**